Parameter study for superradiance and SASE FEL at PITZ

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Estimation of superradiance energy

The spectrum density from a single electron passing an The total energy of the radiation is undulator is given by [1]

$$\frac{d^2 U_m}{d\Omega d\omega} = \frac{e^2 \gamma^2 m^2 K^2}{4\pi\epsilon_0 c(1+K^2/2)^2} \times \frac{\sin^2(\pi N_u(\omega-\omega_m)/\omega_1)}{\sin^2(\pi(\omega-\omega_m)/\omega_1)} \times |JJ|^2$$

$$JJ = J_n \left(\frac{mK^2}{4+2K^2}\right) - J_{n+1} \left(\frac{mK^2}{4+2K^2}\right), m = 2n+1$$

The width of the spectrum is related to the radiation frequency by

$$\frac{\Delta\omega_m}{\omega_m} = \frac{1}{mN_u}$$

The radiation angle is related to the energy of the electron by

$$\Delta\Omega_m = \frac{2\pi}{\gamma^2} \frac{1 + K^2/2}{2mN_u}$$

 $U_m(\omega) = \frac{d^2 U_m}{d\Omega d\omega} \Delta \Omega_m \Delta \omega_m$

For the fundamental mode, the radiation from a bunch of N electrons has the total energy of

$$U_{\omega} = U_0[N + N(N - 1)g^2(\sigma_z)]$$

Assuming a Gaussian distribution, the form factor is given by

$$g^2(\sigma_z) = e^{-4\pi^2 \sigma_z^2/\lambda_s^2}$$

[1] Peter Schmüser, Martin Dohlus, Jörg Rossbach, Chapter 2, Undulator Radiation, **Ultraviolet and Soft X-Ray Free-Electron Lasers, Introduction to Physical Principles, Experimental Results**, Technological Challenges. Springer-Verlag, 2007

Estimation of superradiance energy

• To obtain certain energy, the corresponding bunch length for various charges



Form factor vs bunch length



To get even only 1 nJ at 100 um, the bunch length needs to be ~50 um RMS

DESY.

Estimation of superradiance energy

- Using the shortest RMS bunch length Anusorn has achieved in his simulations, the radiation energy and form factor are calculated
- The radiation energy is **linearly depending** on the bunch charge, because of the very small form factor, thus no coherent radiation; the radiation energy is also quite small



• Besides refining the compression, we may also consider the resonance at a much longer wavelength (maybe sub-THz), in order to get more and measurable radiation

Norm. emittance: 4 um Tran. phase space: matched Long. Dist.: Flattop/Gaussian Correlated energy spread: 0.5% Slice energy spread: 0.05%

Scan parameters:

Bunch charge from 1 nC to 5 nC Peak current from 100 A to 500 A



Longitudinal phase space from S2E simulation with Astra

Norm. emittance: 4 um Tran. phase space: matched Long. Dist.: Flattop/Gaussian Correlated energy spread: 0.5% Slice energy spread: 0.05%

Scan parameters:

Bunch charge from 1 nC to 5 nC Peak current from 100 A to 500 A By Genesis 1.3

- Efficiency or ρ depends on peak current;
 - Radiation energy $\propto e^{-}$ beam energy or charge



Norm. emittance: 4 um Tran. phase space: matched Long. Dist.: Flattop Correlated energy spread: +/-0.5% Slice energy spread: 0.05%

Scan parameters:

Bunch charge from 1 nC to 5 nC Peak current from 100 A to 500 A

A positive energy chirp can almost double the radiation energy







Unit

μJ

μm

μm

ps

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Dynamical effects on superradiant THz emission from an undulator

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For SASE FEL ٠

- Peak current beyond 200-300 A (depending on the bunch charge) doesn't improve the ٠ SASE performance
- The energy chirp is worth to be investigated further to maximum the THz output ٠

Summary

- For superradiance •
 - Shorter bunch length + long / sub-THz radiation wavelength ٠

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INFLUENCE OF AN ENERGY CHIRP ON SASE FEL OPERATION

