

Core + Halo distributions for ASTRA

A procedure to generate core + halo input distributions for ASTRA based on actual laser transverse distribution

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DESY PITZ, Zeuthen, June 2015

Note: MatLAB scripts written by Mikhail Krasilnikov

Overview: Behavior of extracted charge vs laser pulse energy

> Observations:

- The charge extraction is linear for low laser pulse energies (this is how we measure the quantum efficiency, “single particle” regime).
- But for higher laser pulse energies (saturation regime) the charge extraction dependence on laser energy is much weaker (“collective effects” space charge regime).
- We have consistently observed that the extracted charge for high laser pulse energy is larger than that predicted by simulations for a variety of gun settings and laser parameters.
- The extracted charge in the saturation regime depends on specific laser beam parameters and gun operating settings.

> Hypothesis:

- Although the extracted charge saturates in the core of the uniform laser transverse distribution, radial laser halo contributes to additional extracted charge.

> Experiment:

- To test our hypothesis, we have generated initial (input) distributions fitted to first order to the measured laser transverse profiles, which in fact have a radial profile comprised of a flat-top core with Gaussian-like decaying halo.

> Results:

- Using these distributions, we obtain now agreement to first order between ASTRA simulations and measured extracted charge vs laser pulse energy.
- Our observations seem to indicate that halo is contributing to excess extracted charge compared to a uniform core transverse laser distribution.



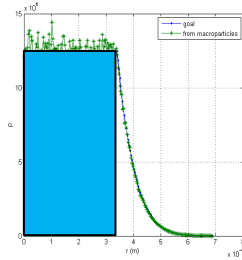
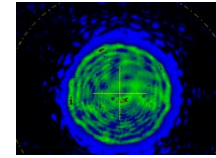
1. Motivation: Using *core+halo* input distributions that match the laser transverse profile agree now with extracted charge measurements
2. Procedure overview
3. Characterization of laser transverse profile
4. Generation of particle input distribution based on laser transverse profile data
5. Procedure summary
6. Sensitivity of simulations and model to laser radial profile parameters
7. Laser radial halo dependence on laser settings and its effect on extracted charge
8. Conclusions



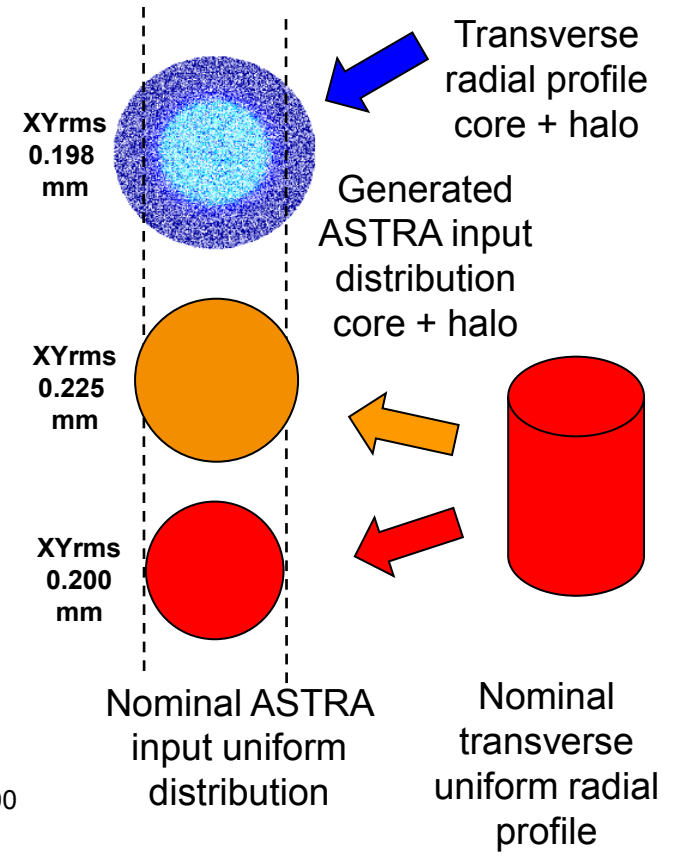
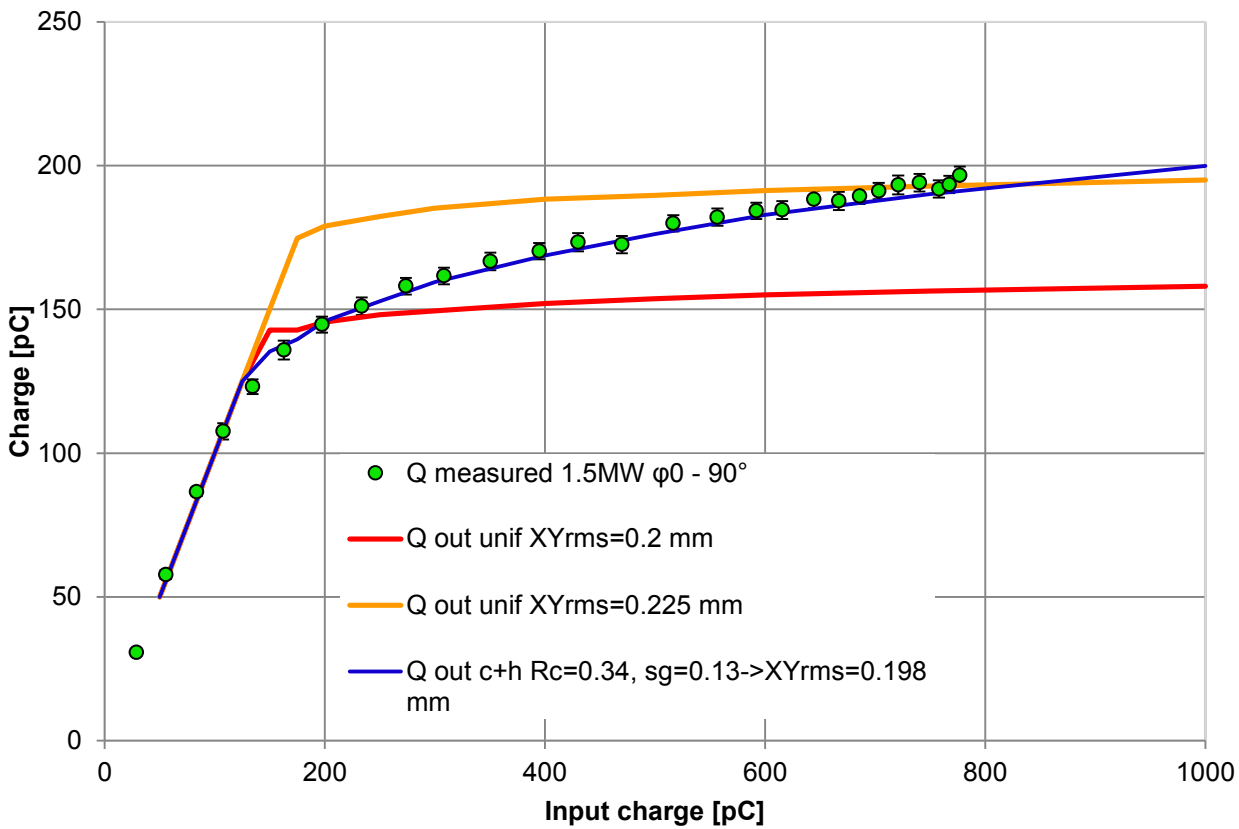
When the core+halo initial distribution is utilized, ASTRA shows good agreement with extracted charge measurements

If a uniform distribution is used instead, the charge saturates

Laser radial distribution image

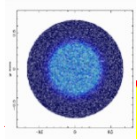
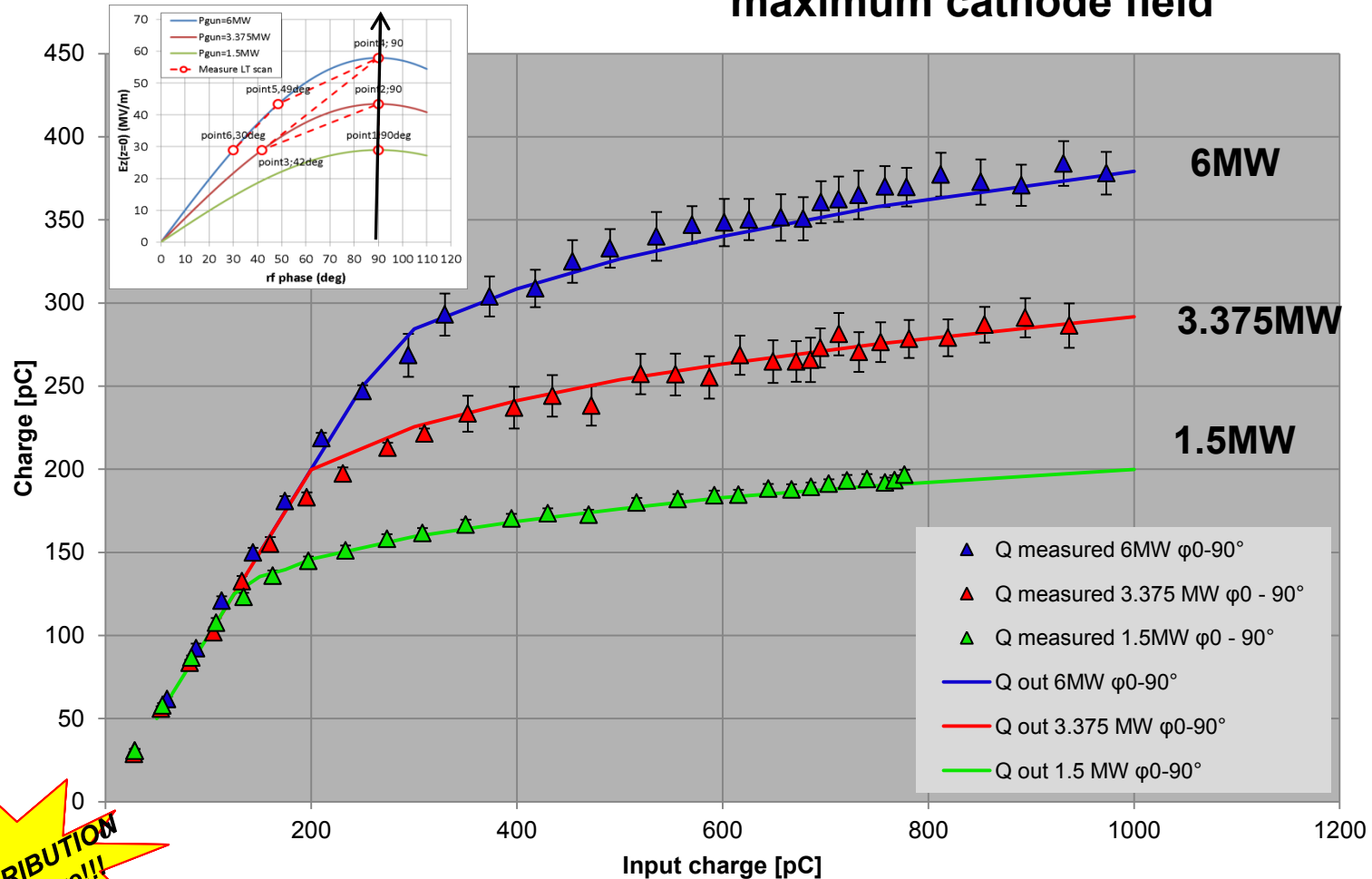


Extracted charge vs laser pulse energy for temporal Gaussian $\sigma_t=1.5$ ps BSA=0.8mm Gun Power = 1.5MW and Gun Phase $\phi_0 - 90^\circ$



Once a fit is found, the core + halo input distribution fits the experimental data...

Extracted charge with core + halo for 0.8 mm beam diameter with 1.5 ps rms Gaussian temporal at maximum cathode field

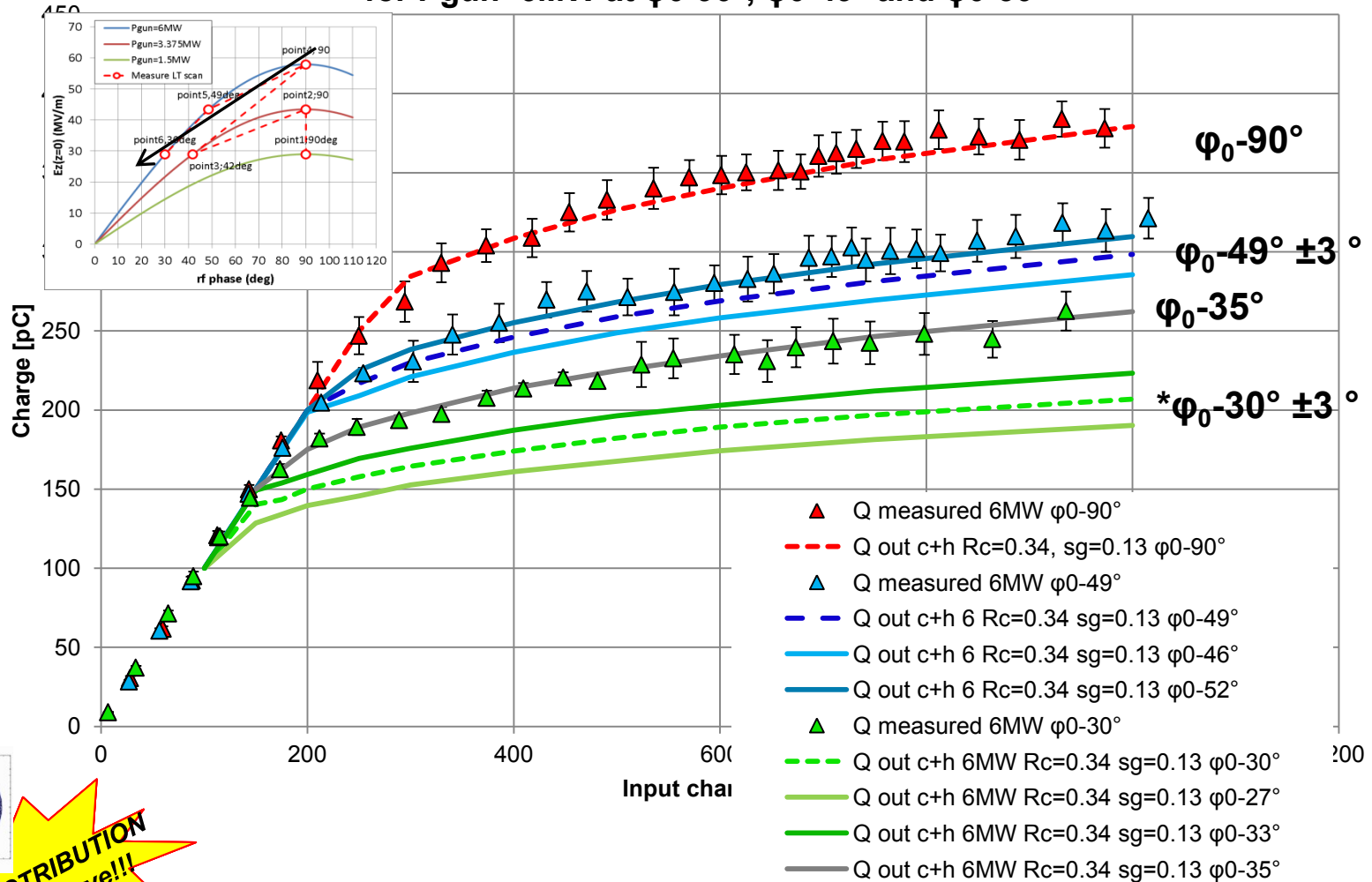


SAME DISTRIBUTION
For each curve!!!

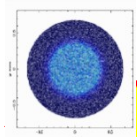


...for most cases....

Extracted charge with core + halo for BSA 0.8 mm with
1.5 ps rms Gaussian temporal
for P_{gun}=6MW at ϕ_0 -90°, ϕ_0 -49° and ϕ_0 -30°



*data taken on different shift, uncertainty in phase is > 5°

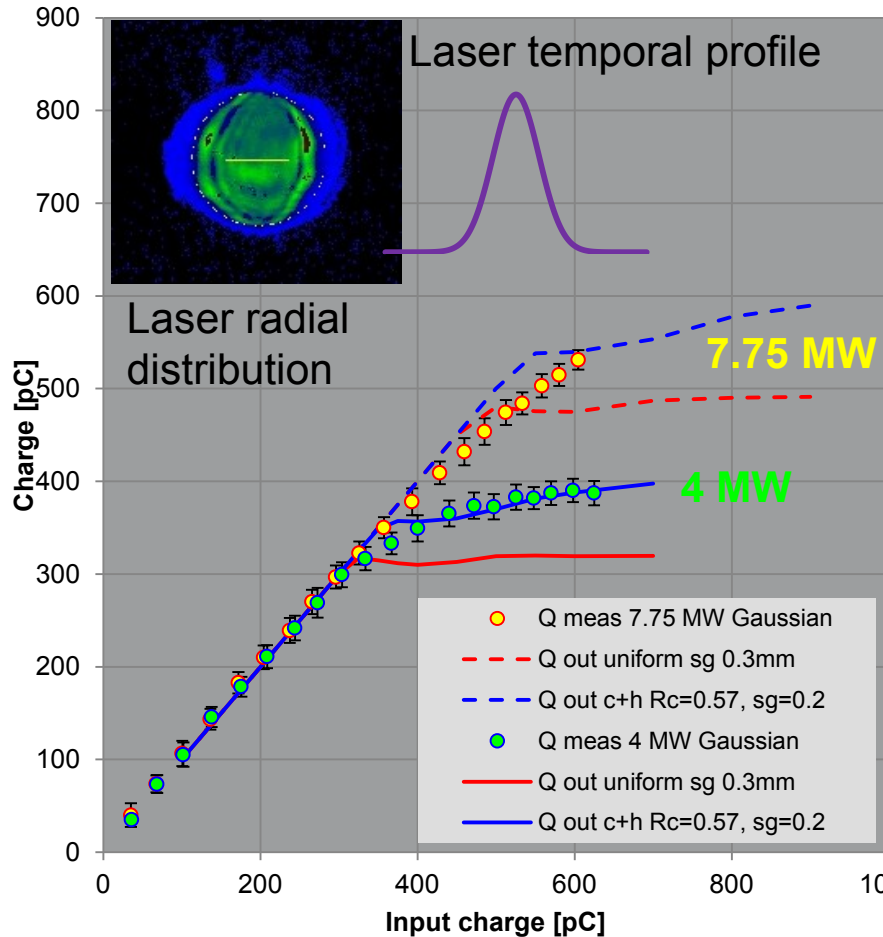


SAME DISTRIBUTION
For each curve!!!

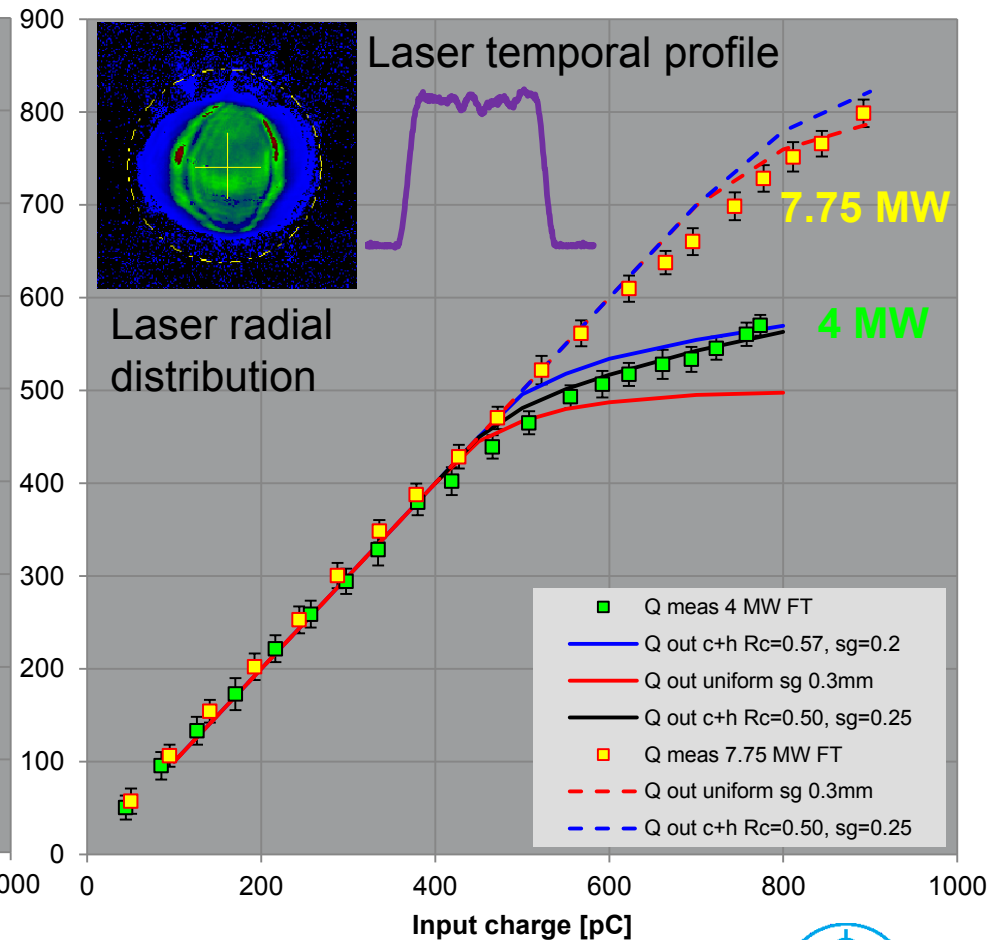


...and to first order for different Gun power, phase, temporal profile and radial size from measurements taken in 2013.

Extracted charge vs laser pulse energy from 2013 for 1.2 mm beam diameter with Gaussian $\sigma_t=1.06\text{ps}$ at MMMG

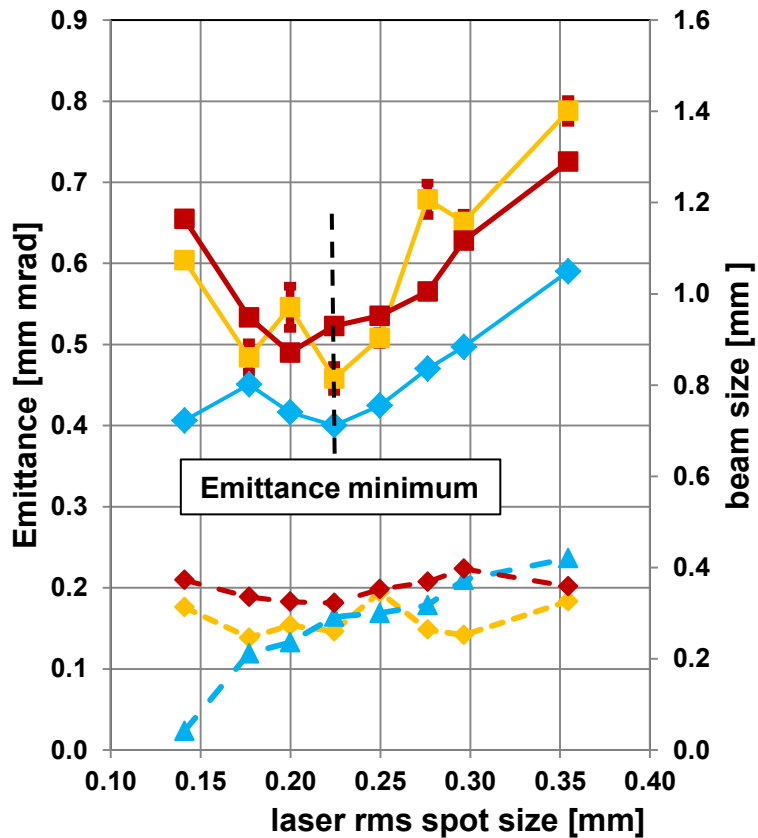


Extracted charge vs laser pulse energy from 2013 for 1.2 mm beam diameter with FWHM=17 ps Flat Top at MMMG

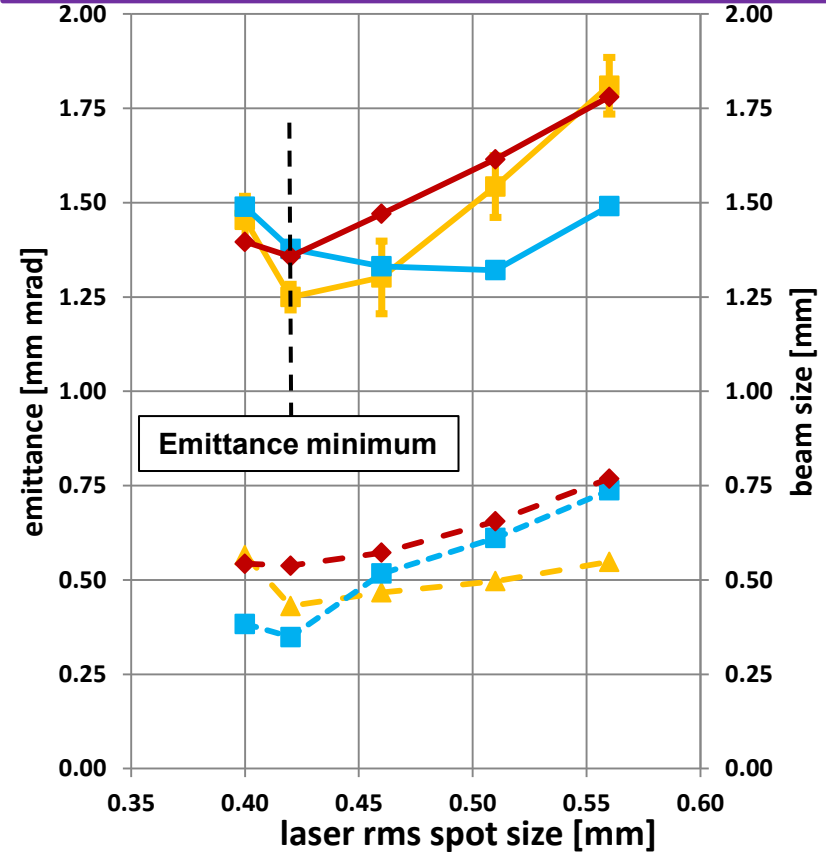


Using core+halo input distributions in ASTRA renders closer agreement with emittance measurements than just using uniform core input distributions*

0.1 nC bunch charge

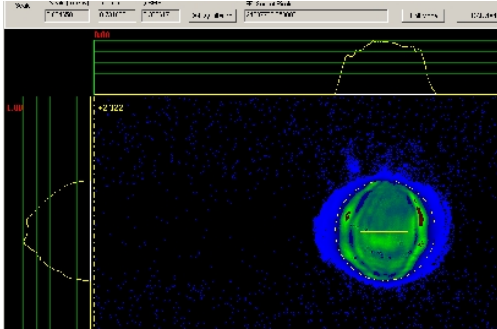


1 nC bunch charge

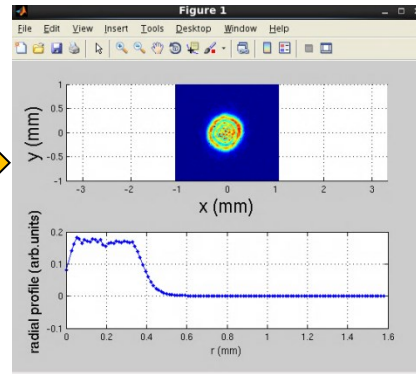


The procedure consists on generating particle input distributions *derived* from the radial curve fit to the measured laser transverse profile:

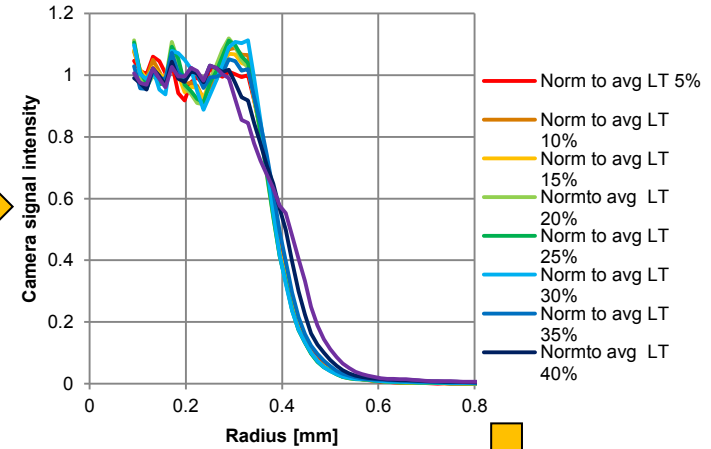
Laser distribution on virtual cathode imaging camera data capture



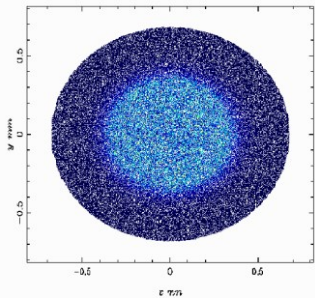
Virtual cathode data fitted in MatLab



Radial laser profile from MatLab data fit

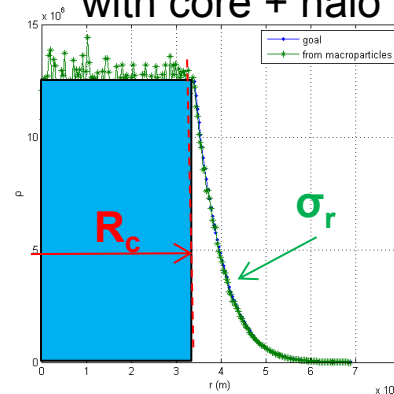


Input distribution shown by postpro with core + halo



To ASTRA

MatLab-generated input distribution for ASTRA with core + halo



Curve fit to match radial laser profile fit

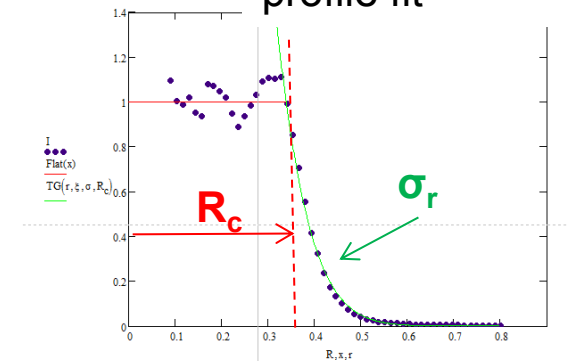
$$TG(r, \xi, \sigma, R_c) := \xi \exp\left(\frac{R_c^2 - r^2}{2\sigma^2}\right)$$

$$\xi := 1$$

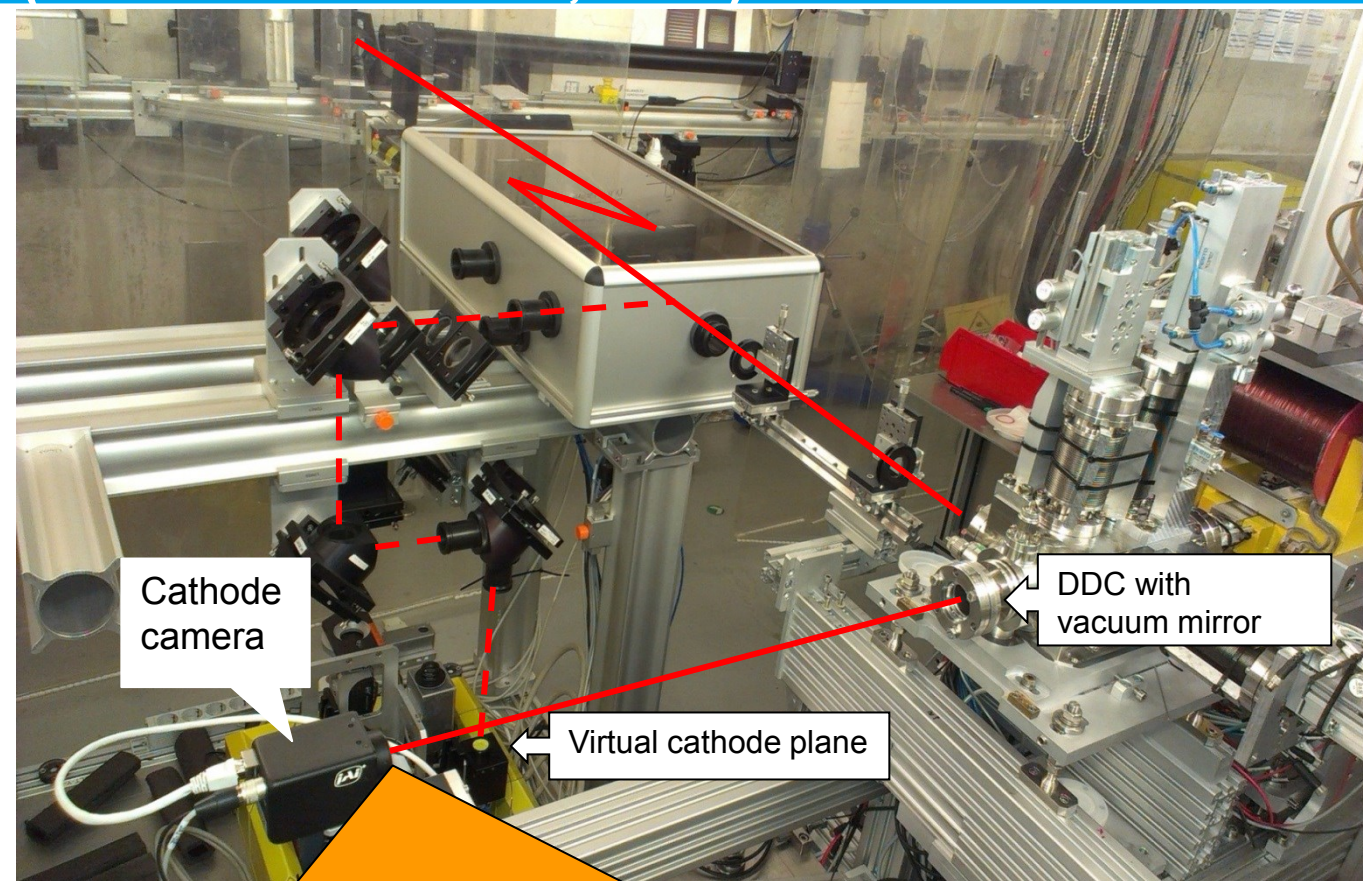
$$R_c := 0.34$$

$$\sigma := 0.15$$

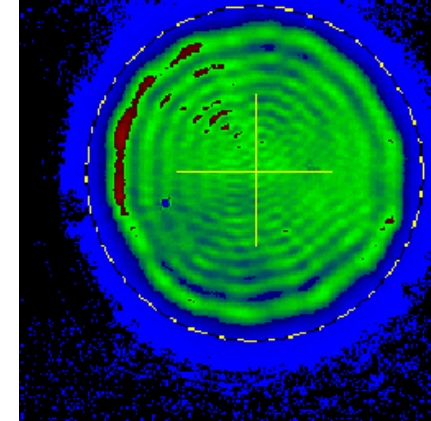
$$r := 0.32, 0.321, \dots, 0.8$$



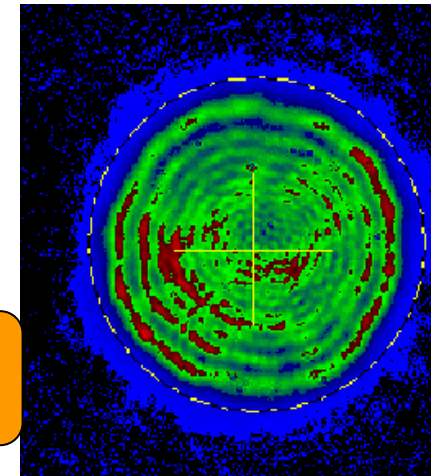
To characterize the laser spot on the cathode: The laser beam shaping aperture (BSA) is imaged onto a CCD camera positioned at a plane equivalent to that of the actual cathode (Virtual Cathode 2, VC2) Data taken 12-March-2013*



Photocathode



Virtual cathode plane

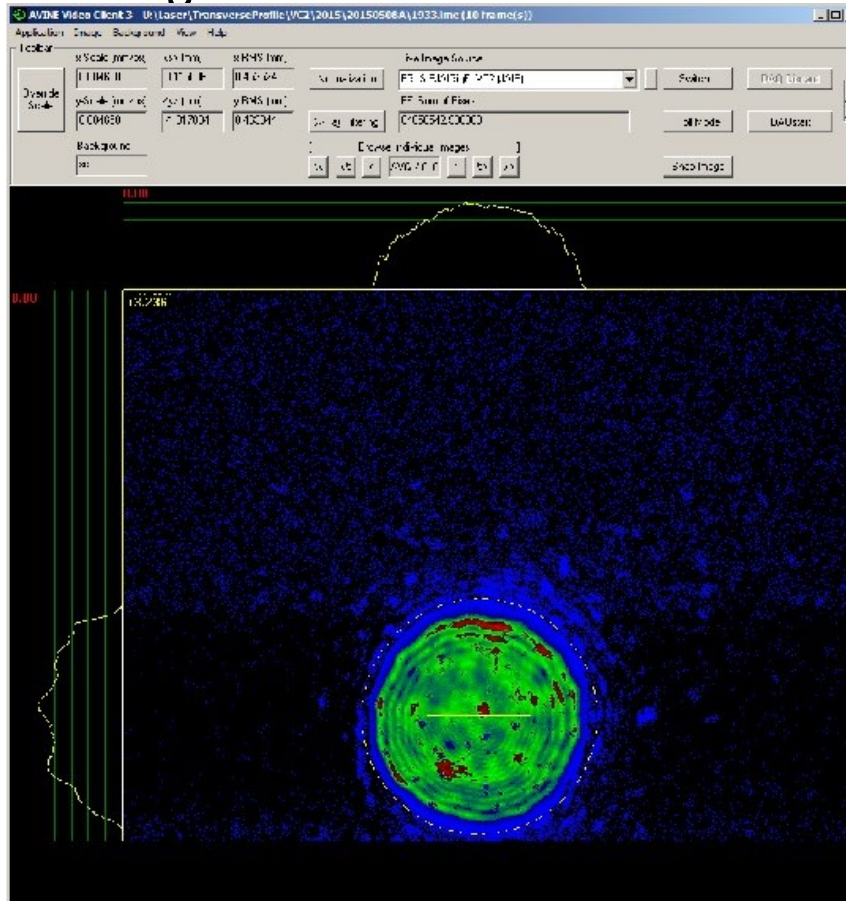


The Cathode camera CCD is placed at the exact location of the Cs2Te photocathode at the back plane of the gun cavity

*M. Gross (PITZ)

Step 1: Capturing VC2 image and data

1. Set BSA to desired value
2. Follow standard procedures to capture VC2 image and data, ensuring the image AND background data are saved in <\\doocs\measure\Laser\TransverseProfile\VC2\YYYY\SHIFT>
3. Record VC2 image from Avine video client for reference, as shown below:



➤ Typically 2 x 20 frames (background + signal)

Step 2: Reading the saved VC2 image & background data

1. Open the MatLab script “*read_imc.m*” in LINUX environment. This script is saved in
2. Edit the script to set the **folder** from which the image and background **data file** will be loaded, as shown below:

```
clear all;
close all;
[imc_data, header_data] = avine_load_video_images_from_file('/doocs/measure/Laser/TransverseProfile/VC2/2015/20150514N/0027_LT16_BSA2.imc');
[bkc_data, header_data_bkc] = avine_load_video_images_from_file('/doocs/measure/Laser/TransverseProfile/VC2/2015/20150514N/0027_LT16_BSA2.bkc');
header_data_single = header_data(1);
camWidth = header_data_single{1}.width;
camHeight = header_data_single{1}.height;

scale = 0.00465; % mm/pixel

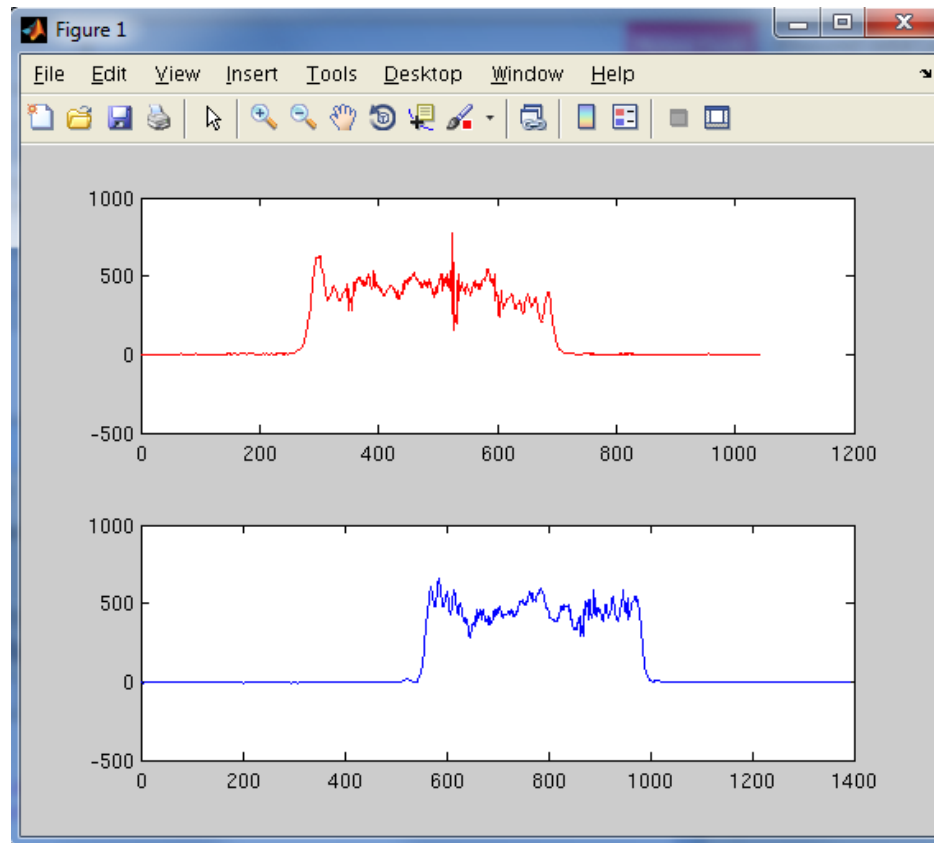
averageVector = zeros(camHeight,camWidth);
averageVectorBkc = zeros(camHeight,camWidth);
```

3. Run the script



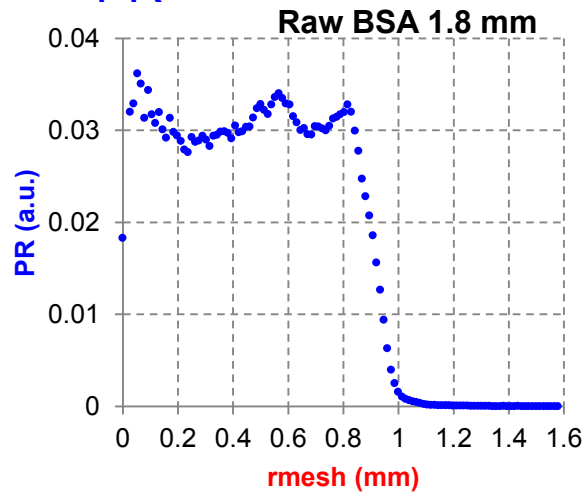
Step 2 Continued

- The Script only reads the image data and subtracts the background.
- The only output is the display of the vertical and horizontal cuts, just to indicated it has finished reading the data.



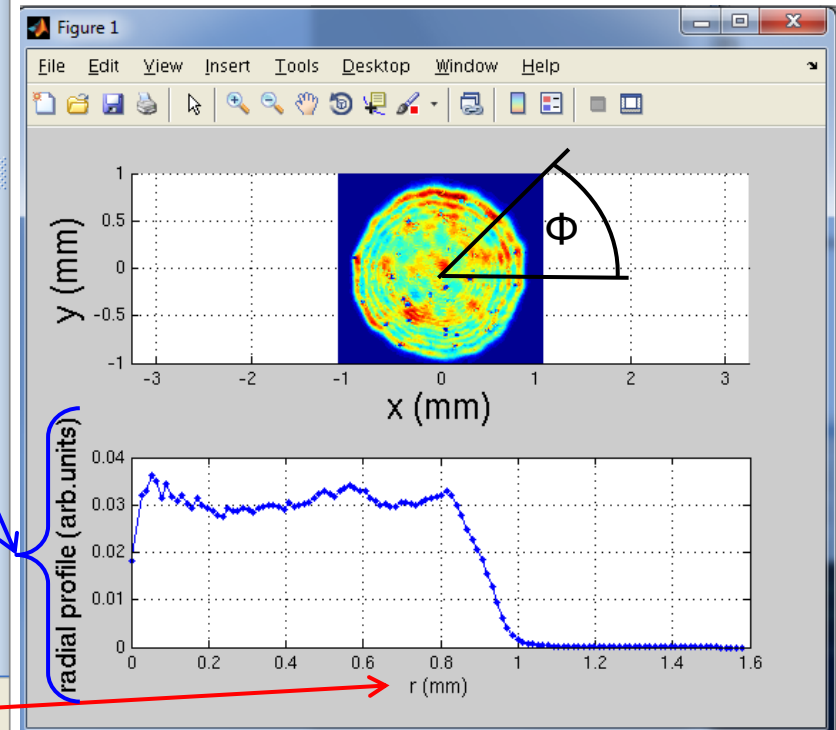
Step 3: Generating radial profile averaged over Φ

1. Open the MatLab script "*RadProfiler.m*" in LINUX environment. This script is saved in
2. Run the script
3. Make a radial profile in Excel or any other graphing tool with data stored in **rmesh** and **PR**



Name	Value
A	<461x461 do
C	<2x151188 d
CF	<2x153444 d
CenterBin	1.7608e-04
CutX	<1040x1 dou
CutY	<1x1392 dou
lcur	-1.5128e-08
M	461
N	461
Nr	119
PR	<1x120 doub
R	<461x461 do
Rmax	1.5783
X	<461x461 do
Y	<461x461 do
ans	'PLOTTING...'
averageVector	<1040x1392
averageVector...	<1040x1392
bkc_data	<50x1 cell>
bkc_data_single	<1x1 cell>
bkc_vector	<1040x1392
camHeight	1040
camWidth	1392
count	21
del	230
dr	0.0132
h	<8050x1 dou
header_data	<50x1 cell>
header_data_bkc	<50x1 cell>
header_data_si...	<1x1 cell>
hf	1
i	461
imc_bkc	<1040x1392
imc_data	<50x1 cell>
imc_data_single	<1x1 cell>
imc_vector	<1040x1392
ir	114
j	461
numFrames	50
numFramesTe...	[50,1]
r0	0.0263
rcur	1.5125
rmesh	<1x120 doub
s0	6.2135e+07
sc	0.0047

- > The script takes the matrix data generated by *read_imc.m* and averages over Φ .
- > It generates the graph and radial profile shown below
- > The radial profile data is saved in the vectors **rmesh** and **PR**.



Step 4: Finding core + halo parameters R_c and σ_r from radial profile

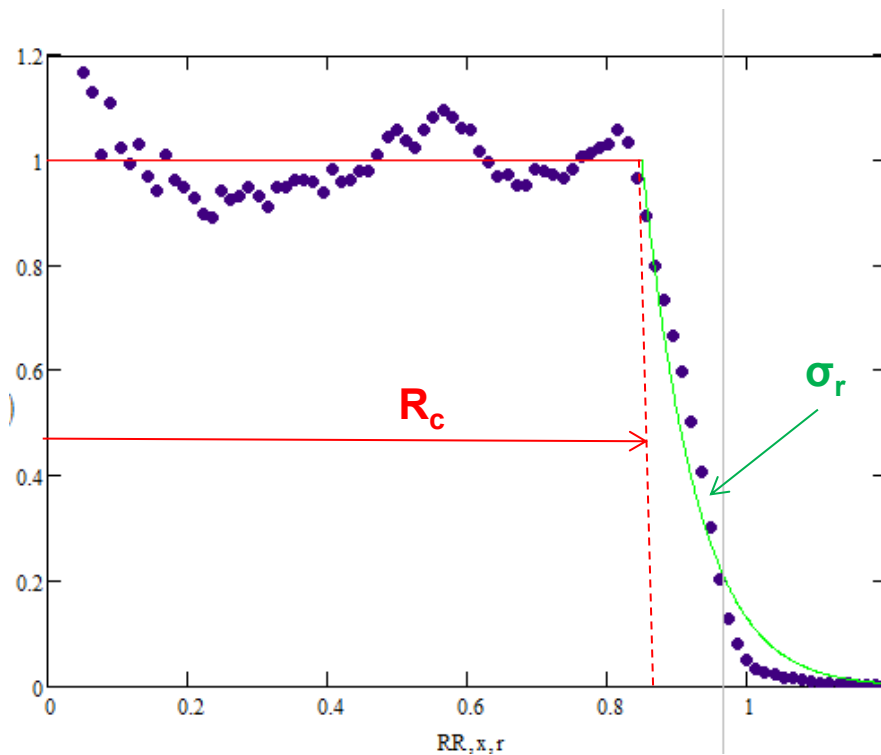
1. From the radial profile, find R_c and σ_r by fitting the following equation to the radial profile:

2.
$$F(r) = \exp\left[\frac{R_c^2 - r^2}{2\sigma_r^2}\right]$$

> This can be accomplished manually with MathCAD for example, or a MatLab script could be written for this purpose.

> For this particular case:

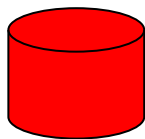
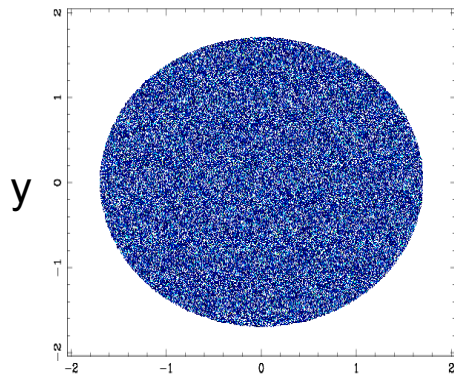
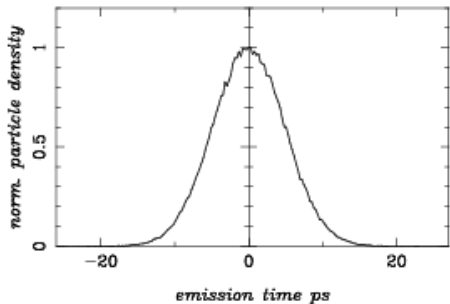
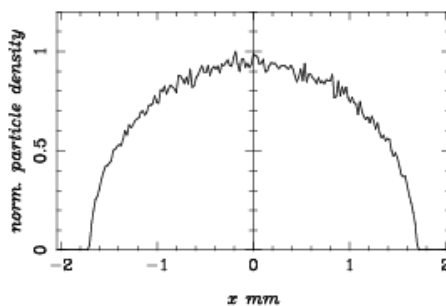
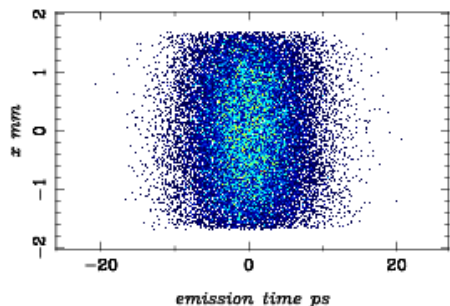
- $R_c = 0.85 \text{ mm}$
- $\sigma_r = 0.26 \text{ mm}$



Step 5: Generating baseline uniform particle distribution needed for core + halo distribution

1. Edit a generic **generator.in** file to create a **uniform** distribution with:

- Desired **temporal profile**, and
- RMS beam size must be equal to Rc**



C. Hernandez-Garcia & M.

```
gen_unif_for_c+h_BSA_1p8mm.in - WordPad
| &INPUT
  FNAME = 'for_c+h_BSA_1p8mm.ini'
  ADD=.F
  N_add=0
  IPart=200000
  Species='electrons'
  Probe=.True.
  Noise_reduc=.F.
  High_res=.T.
  Cathode=.T.
  Q_total=0.5
  Ref_zpos=0.0E0
  Ref_clock=0.0E0
  Ref_Ekin=0.0E0
  Dist_z='gauss',
  sig_clock=4.8936E-3,
  Dist_pz='i',
  sig_Ekin=0.0E0,
  emit_z=0.00E0,
  cor_Ekin=0.E0,
  LE=0.00055
  Dist_x='r',
  sig_x=0.85
  Dist_y='r',
  sig_y=0.85
  Dist_px='r',
  Nemit_x=0.0E0,
  cor_px=0.0E0
  Dist_py='r',
  Nemit_y=0.0E0,
  cor_py=0.0E0
  LPROMPT=.F
/
```

Also Flat-Top
If necessary

$E_k = 0.55$ eV
thermal emittance

Step 6: Generating core + halo distribution with R_c and σ_r based on uniform distribution

1. Open the MatLab script “*core_halo generator for ASTRA.m*”, saved in

....

2. Edit the script to:

1. Load the uniform baseline distribution

2. Set R_c

3. Set σ_r

4. Save the core + halo distribution

> NOTE: The script needs the macro named: “*Rdistr.m*”

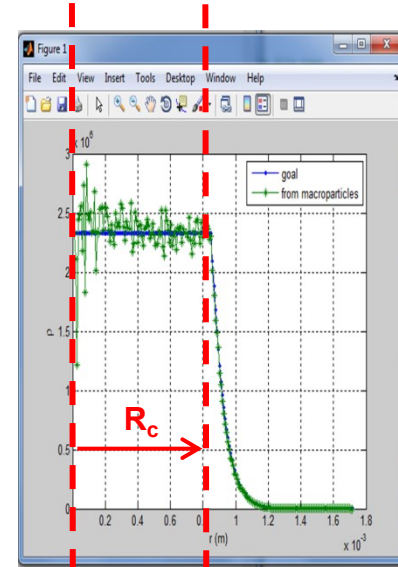
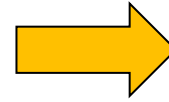
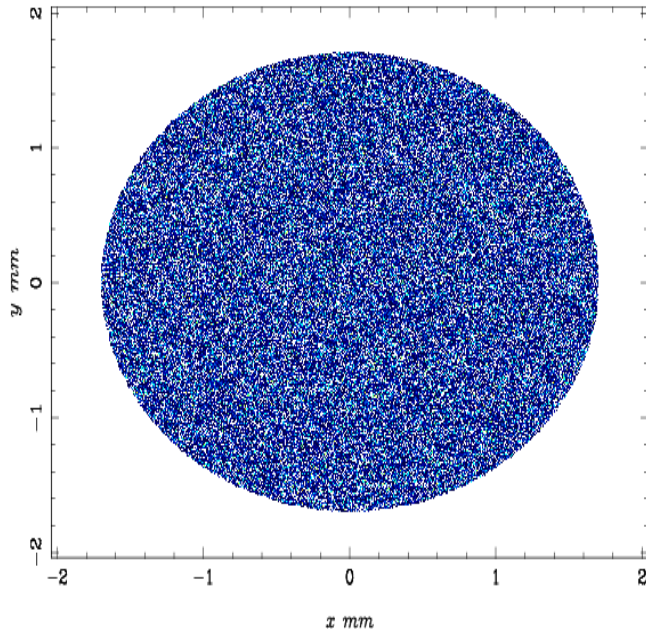
```
Editor - \\win.desy.de\home\carlos\My Documents\PITZ\Emission model dev...
1 - clear;
2 - clc;
3
4 - fn='for_c+h_BSA_1p8mm.ini'; %nonperturbed dist
5 - %fn='test10k.ini'
6
7 %ETHA=1.5 % signaHalo/sigmaCore
8 - xi=1.0 %This defined the height of the Gaussia
9
10 %Based on manual fit for BSA 0.8mm
11 - Rc=0.85;
12
13 %from here everything in meters!
14 - Rc=Rc*1e-3; % mm -> m
15
16 - sigH=0.26;
17 - sigH=sigH*1e-3;
18 - drg=Rc/100;%grid for the profile
19
20
21 - rg=0.0*drg:drg:Rc*2+2*drg;
22 - Ngrid=max(size(rg)) ;
23 - RHOgoal=rg.*0;
24 - PDF=RHOgoal;
25
26 - for n=1:Ngrid
27 -     RHOgoal(n)=Rdistr(rg(n), Rc, xi, sigH);
28 - end
29
30
31
32
33 - fno=['c+h_BSA_1p8mm']; %output file name
34 - fno=strrep(fno, '.', 'p');
35 - fno=[fno, '.ini'] %output ASTRA ini file
36 - fid=fopen(fno, 'wt');
```

Step 6: This is the last step in the procedure

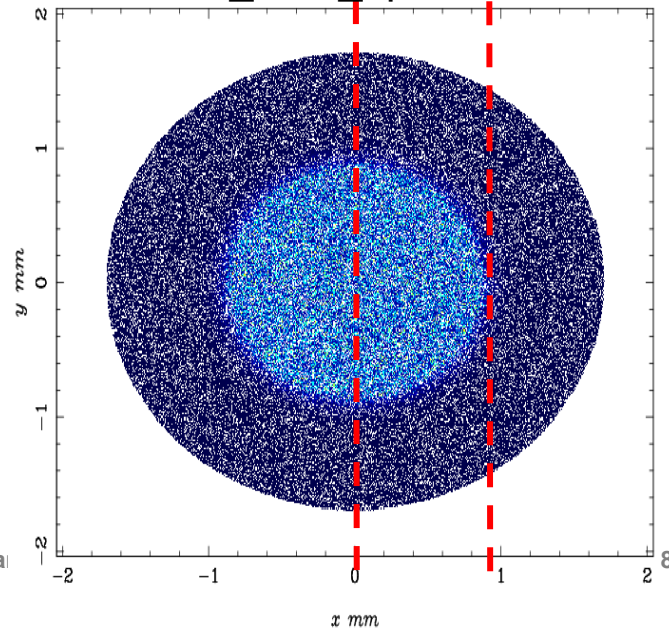
- The script takes the uniform baseline distribution and makes a new distribution composed of a flat-top core with radius R_c , and a decaying Gaussian-like halo shifted by R_c and with σ_r by modification of the macroparticle charge



for_c+h_BSA_1p8mm.ini



c+h_BSA_1p8mm.ini



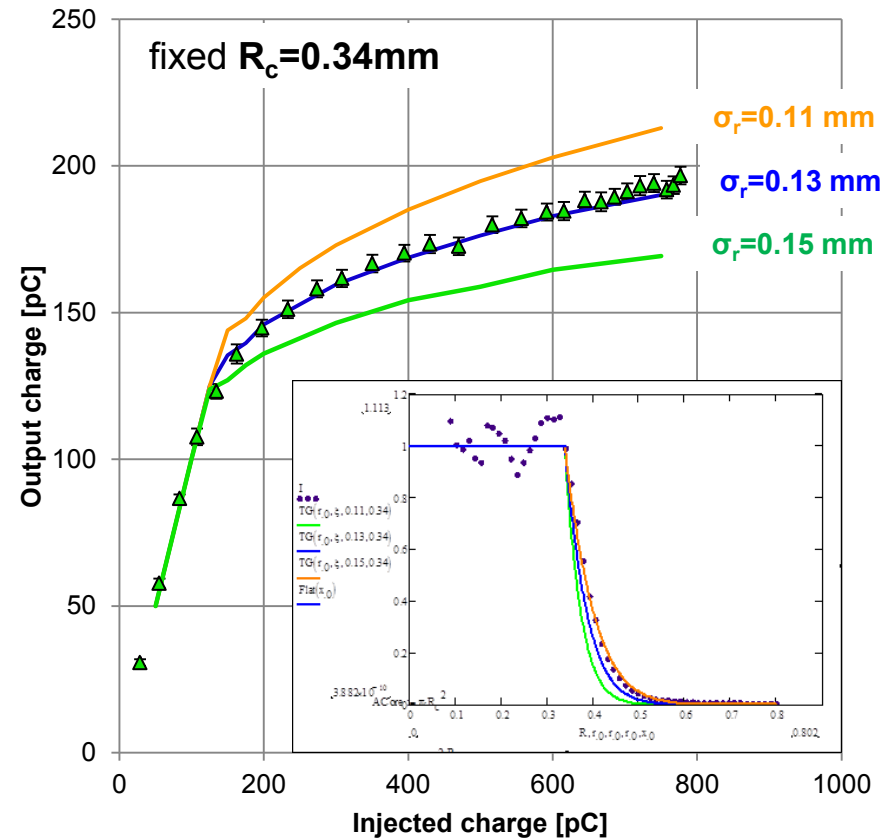
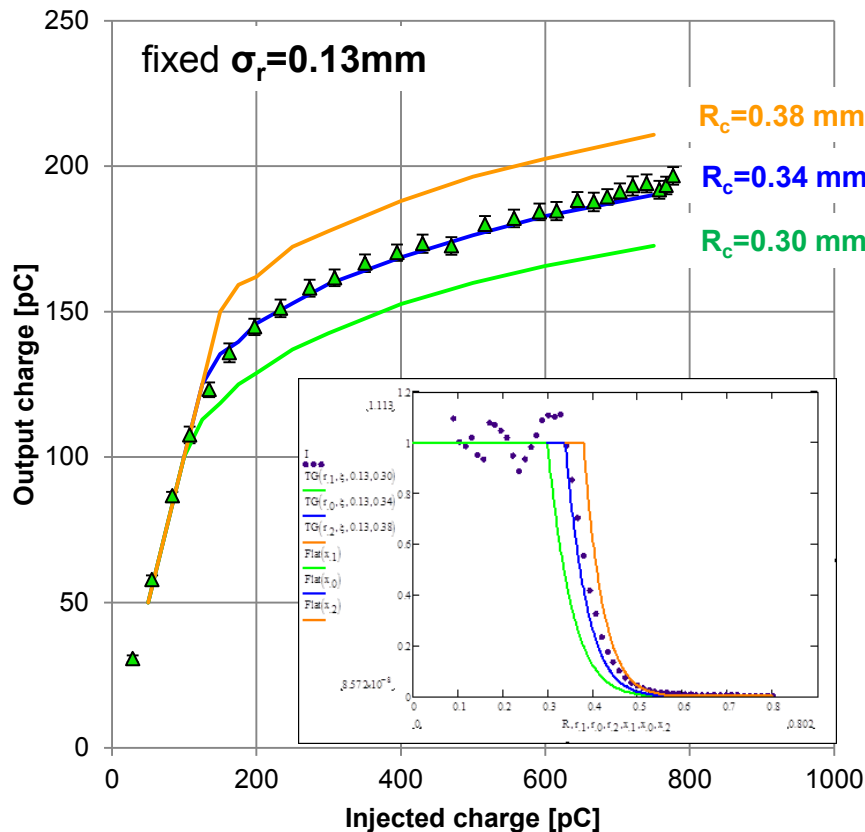
Procedure summary

1. Capture VC2 image and data
2. Run MatLab script “read_imc.m” to read VC2 data
3. Run MatLab script “*RadProfiler.m*” to generate radial profile
4. Find core + halo parameters (R_c , σ_r) from the radial profile
5. Generate baseline uniform particle distribution with $XY_{rms} = R_c$
6. Run MatLab script “*core_halo generator for ASTRA.m*” with R_c , σ_r and baseline uniform distribution to generate core + halo input distribution to be used in ASTRA

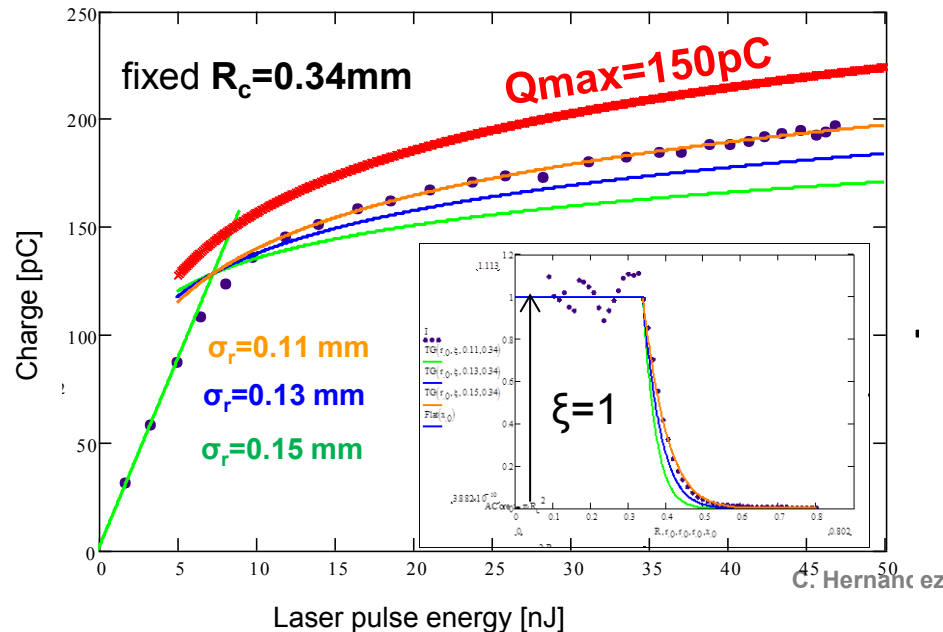
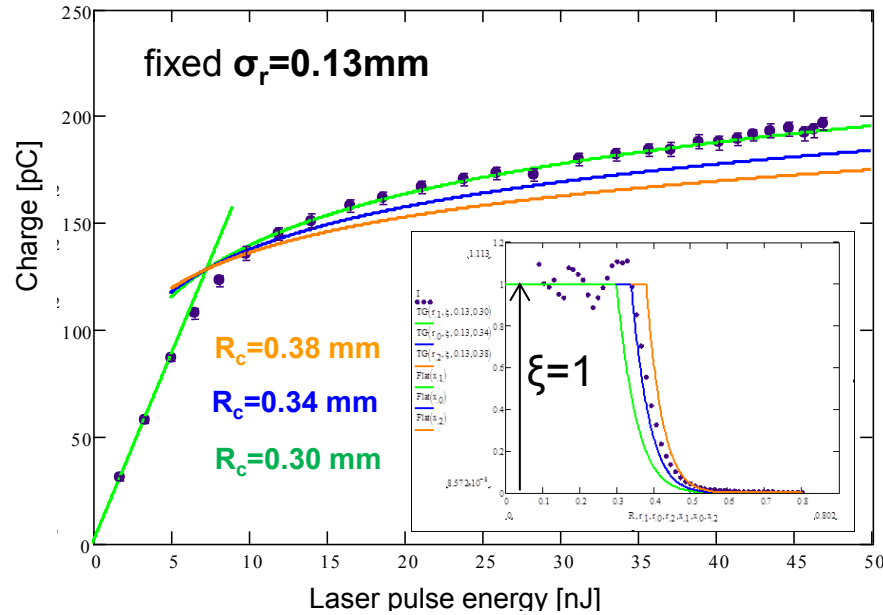


The following graphs show the sensitivity of ASTRA simulations to the core + halo parameters R_c and σ_r

Extracted charge vs laser pulse energy for temporal Gaussian $\sigma_t=1.5$ ps
BSA=0.8mm Gun Power = 1.5MW and Gun Phase $\phi_0 - 90^\circ$



The following graphs shows the sensitivity of the semi-analytical model to the core + halo parameters R_c and σ_r

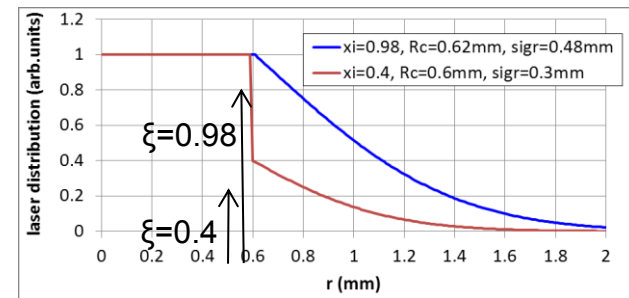


Model by Mikhail Krasilnikov (PITZ)

$$F_l(r) = \frac{E_l}{\pi R_c^2 + 2\pi\xi\sigma_r^2} \begin{cases} 1, & \text{if } r \leq R_c \\ \frac{R_c^2 - r^2}{\xi e 2\sigma_r^2}, & \text{if } r > R_c \end{cases}$$

$$\eta = \frac{2\sigma^2}{R^2}$$

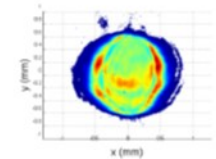
$$\rho_Q(r) = \frac{2QE \cdot E_l}{\pi R_c^2 + 2\pi\xi\sigma_r^2} \begin{cases} 1, & \text{if } r \leq R_c \\ \frac{R_c^2 - r^2}{\xi e 2\sigma_r^2}, & \text{if } r > R_c \end{cases}$$



$$Q = Q_{core} + Q_{halo}$$

$$Q_{core} = \frac{1}{1 + \xi \cdot \eta} \begin{cases} Q_{exp}, & \text{if } Q_{exp} \leq Q_{max} \\ Q_{max}, & \text{if } Q_{exp} > Q_{max} \end{cases}$$

$$Q_{halo} = \frac{\eta}{1 + \xi \cdot \eta} \begin{cases} \xi \cdot Q_{exp}, & \text{if } \xi \cdot Q_{exp} \leq Q_{max} \\ Q_{max} \cdot \left(1 + \ln \frac{\xi \cdot Q_{exp}}{Q_{max}}\right), & \text{if } \xi \cdot Q_{exp} > Q_{max} \end{cases}$$

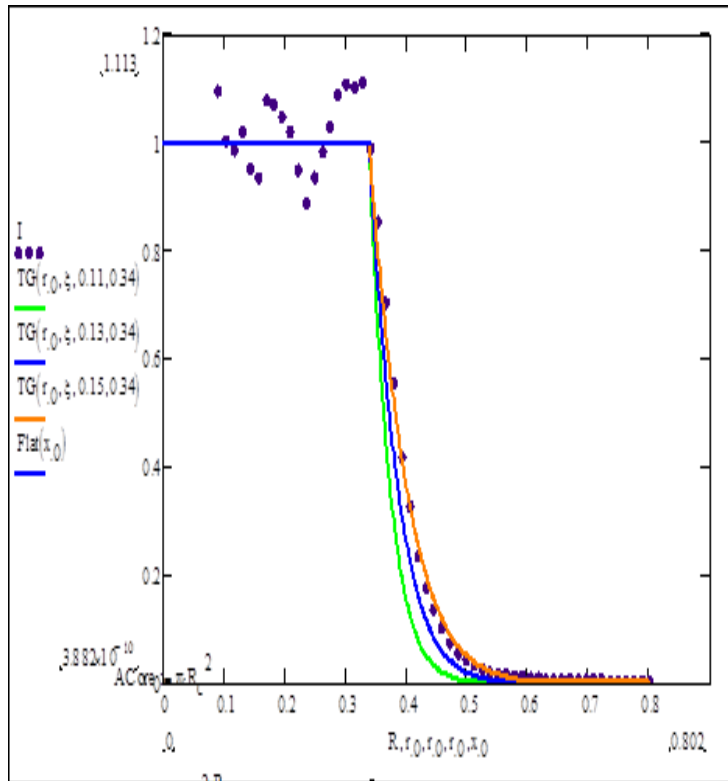


Q_{max} is the only fit parameter (weight factor) in the semi-analytical model. $Q_{max} = 128 \text{ pC}$ except where indicated

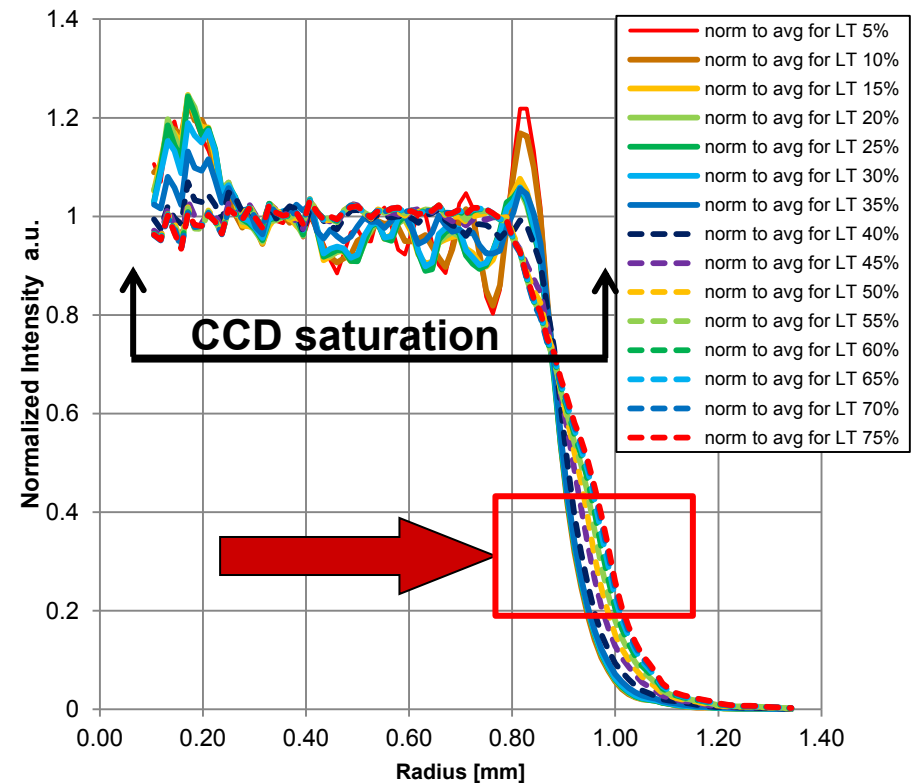


CAUTION!!! Note that halo changes also with increased Laser Transmission (LT).

- > This means that if the VC2 image is captured with low LT, R_c and σ_r are found fitting that laser radial profile



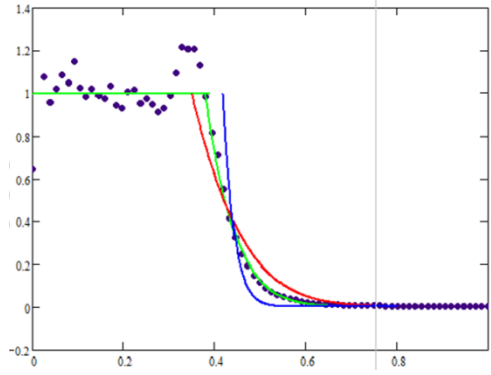
- > HOWEVER often the emittance is measured with higher LT values, which are not VC2 captured, but that certainly changes the halo profile



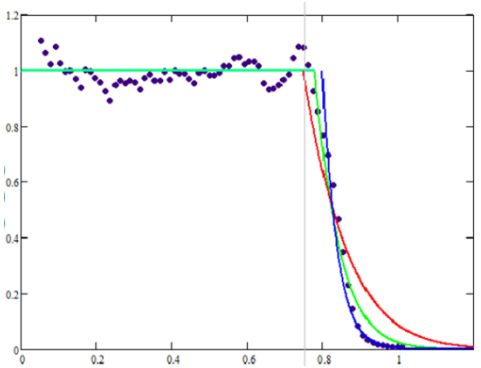
Laser halo intensity distribution effect to emittance (Q. Zhao)

➤ The laser on VC2(virtual cathode) data recorded with low Laser transmission (LT)~18%, but for high bunch Charge, the LT will be much higher. So from above simulation, the halo fit from data On VC2 is not considered the halo intensity distribution.

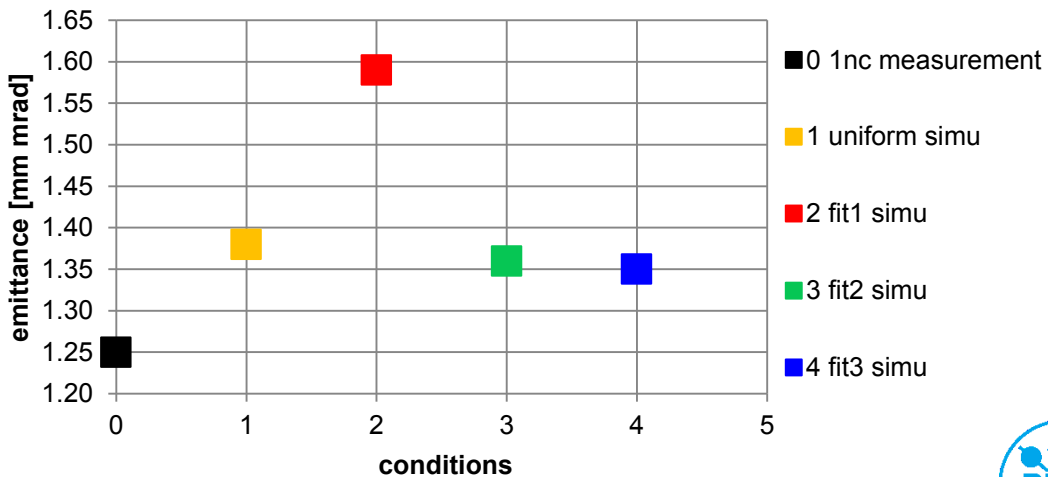
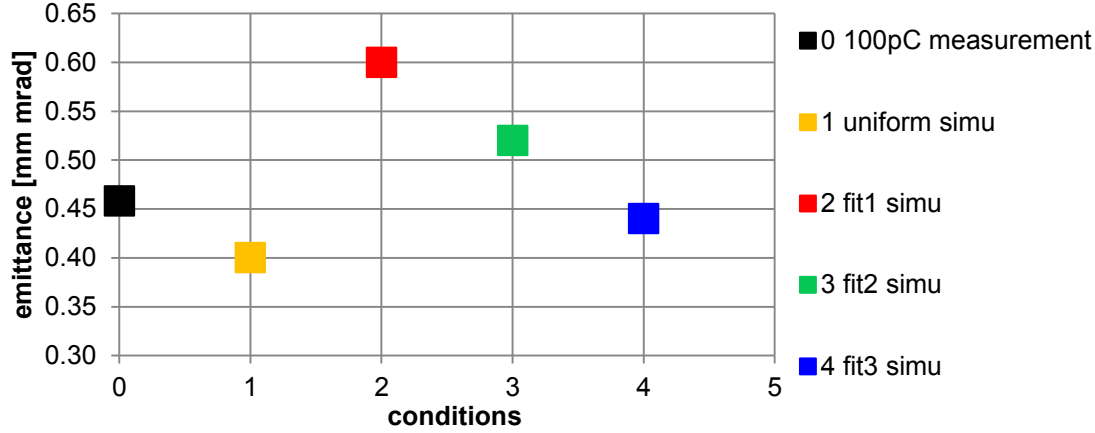
100pC BSA= 0.9mm



1 nC BSA= 1.6mm



conditions: 0-measurement, 1-uniform simu, 2-fit1 simu
3-fit2 simu, 4-fit3 simu



The imaged laser spot on the virtual cathode plane shows smaller halo / core ratio for larger beam diameters...

0.8 mm

1.0 mm

1.4 mm

1.8 mm

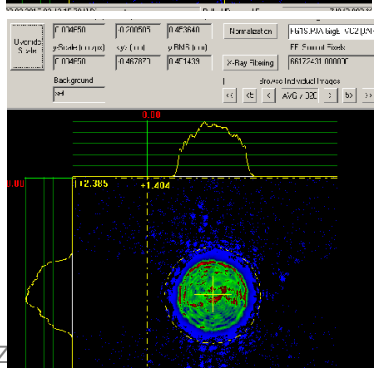
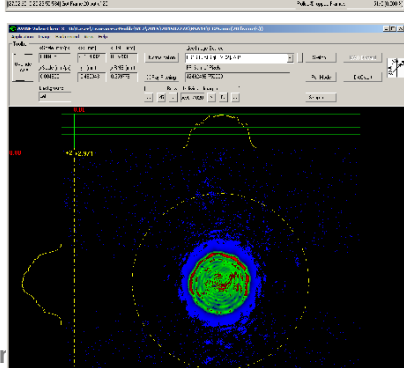
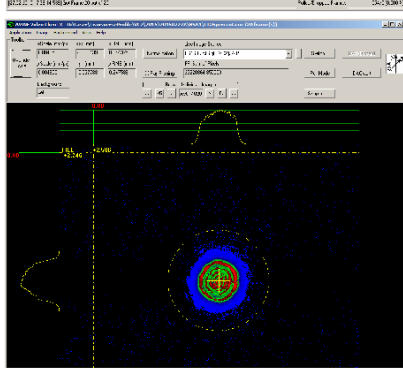
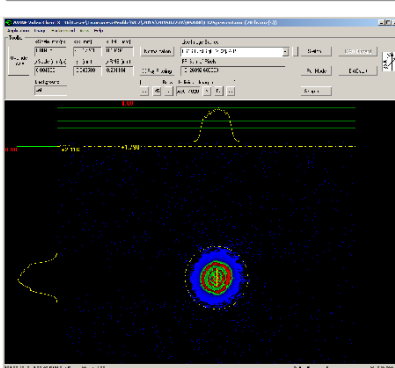
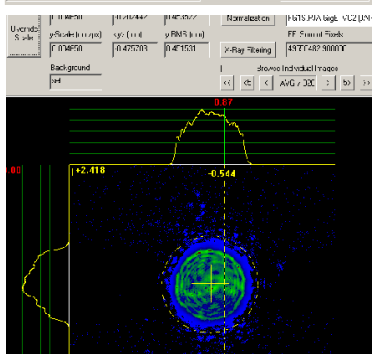
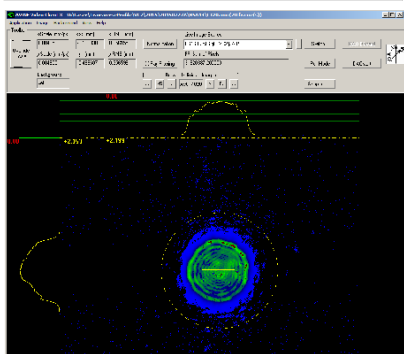
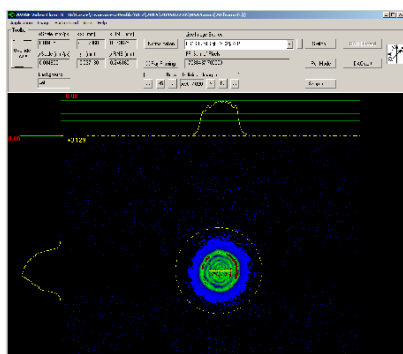
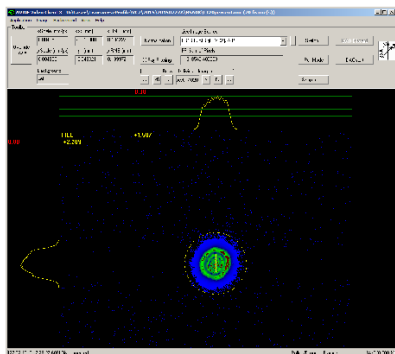
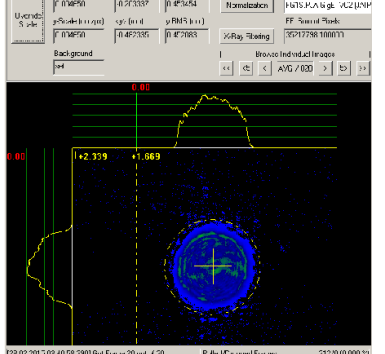
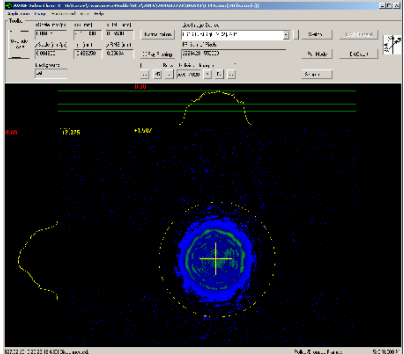
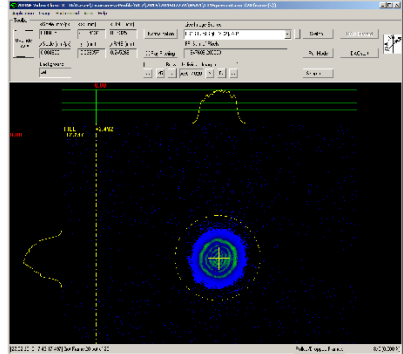
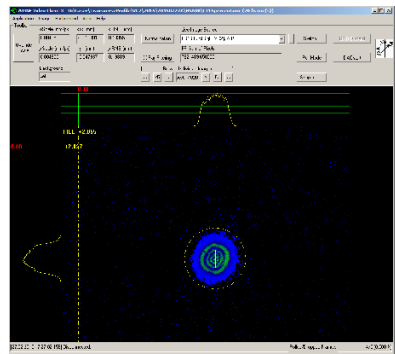
← Beam diameter

Relative Laser Intensity

15%

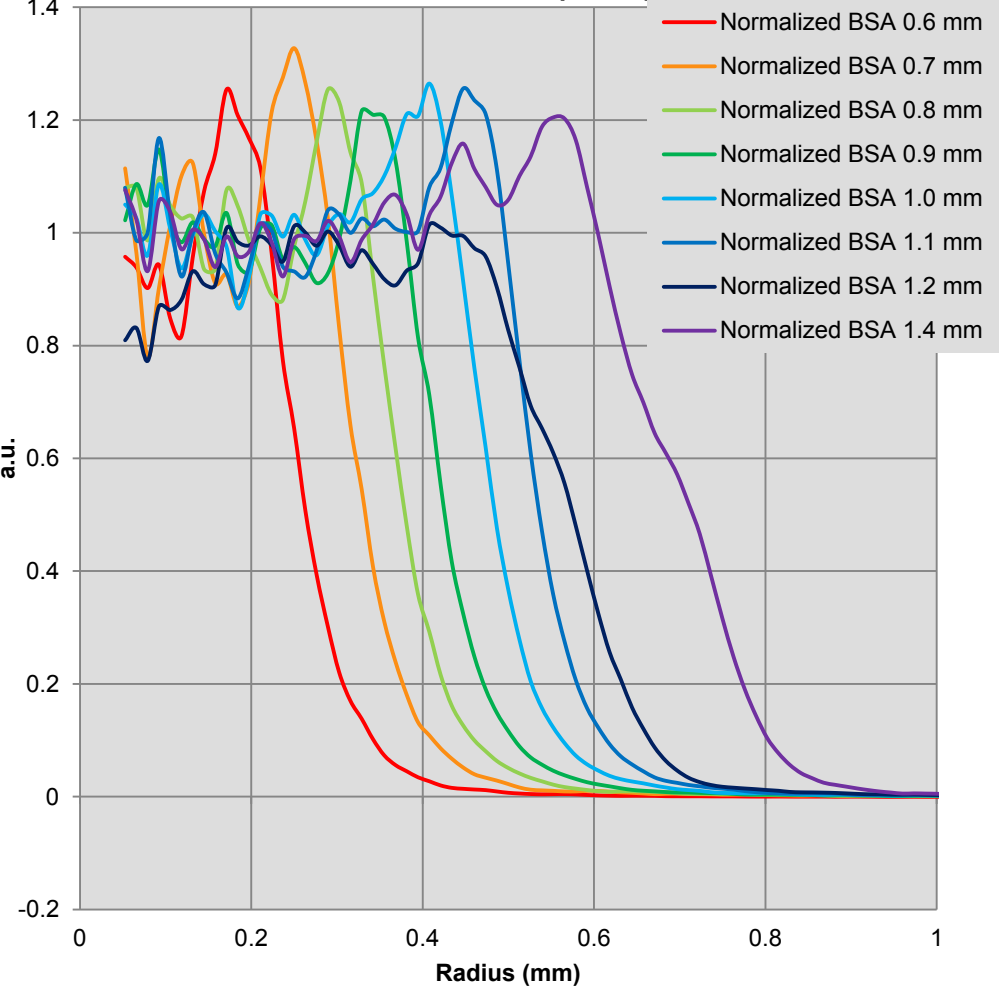
20%

25%

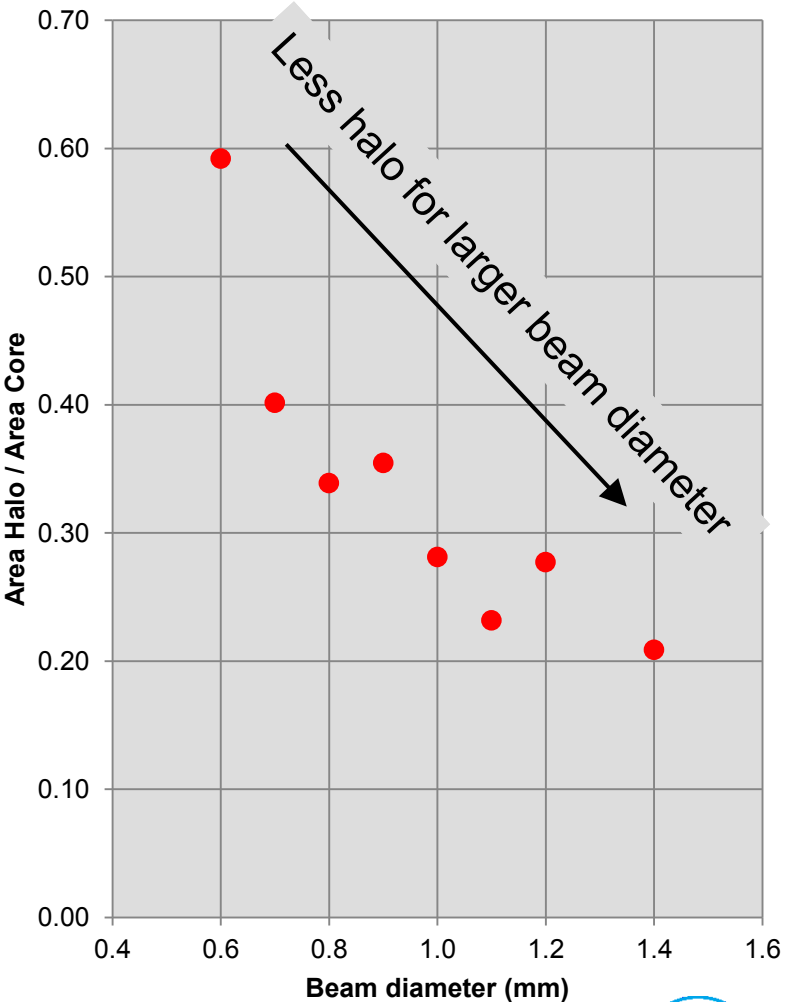


...this means there is less Halo for larger BSA settings...

Laser radial profiles extracted from the cathode imaging plane for various beam diameters (BSA)



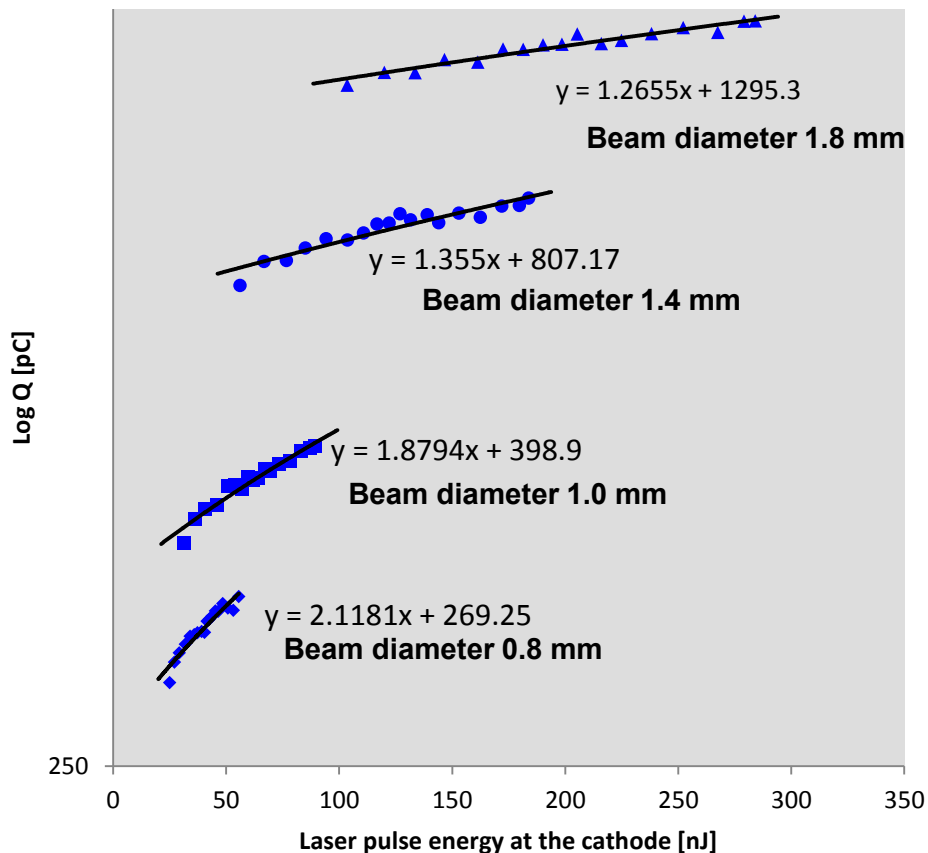
Ratio of Area_Halo / Area_Core



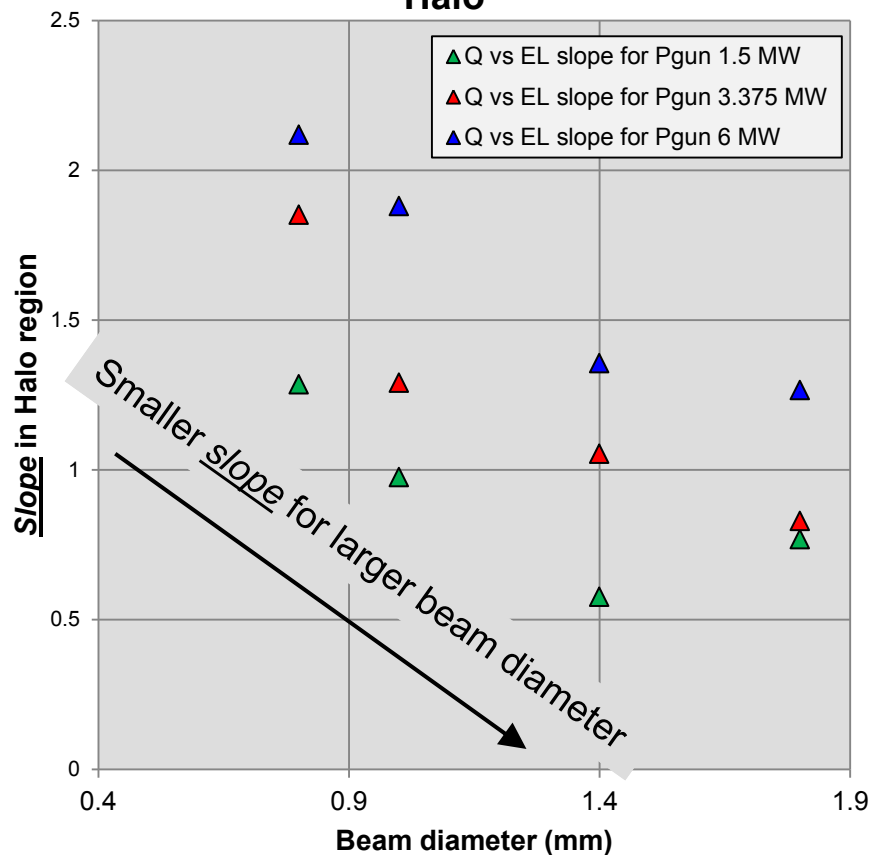
...which is consistent with measurements showing that less charge is extracted from Halo for larger BSA settings

- The slope indicates that charge continues to be extracted from the saturated regions even though charge from core has saturated

Example: Pgun 6 MW

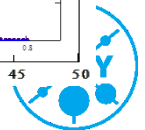
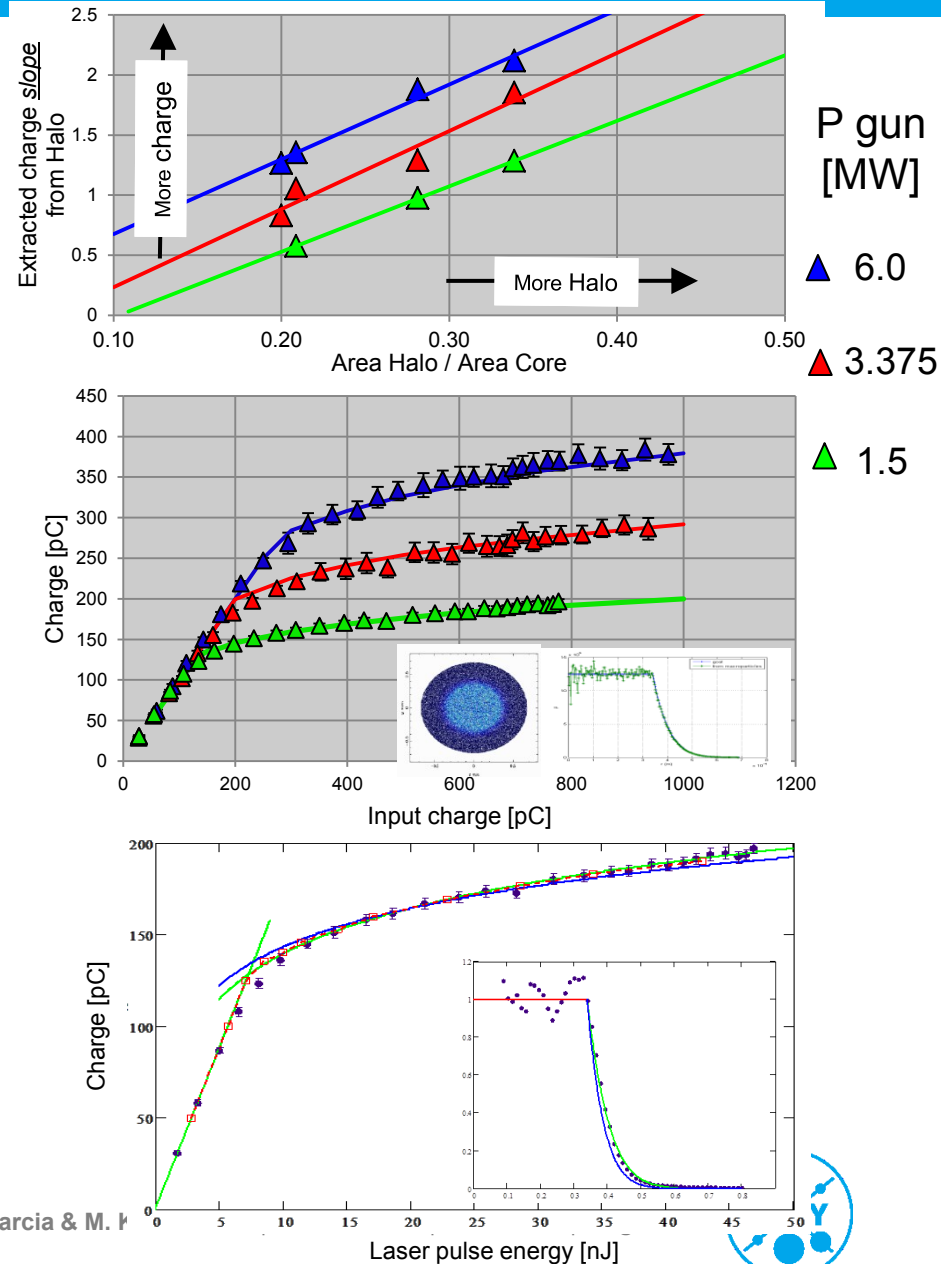


Q vs Laser pulse energy SLOPE: Large slope = more charge -> more Halo



Conclusions:

- > The relationship between the **amount** of **halo** in the measured **laser** radial distribution seems to be **proportional** to the **amount** of **extracted charge** in the saturated emission region.
- > The **measured** charge vs laser pulse energy can be reproduced by **ASTRA** simulations when **core + halo** radial profiles are utilized as input distributions based on fits to actual laser radial profiles.
- > Using M. Krasilnikov's **model** with the **core** radius and **halo** Gaussian σ_r found from the laser radial profile, and fitting only Q_{max} , one gets fairly good agreement with **ASTRA** and with experimental **measurements** in the saturated region

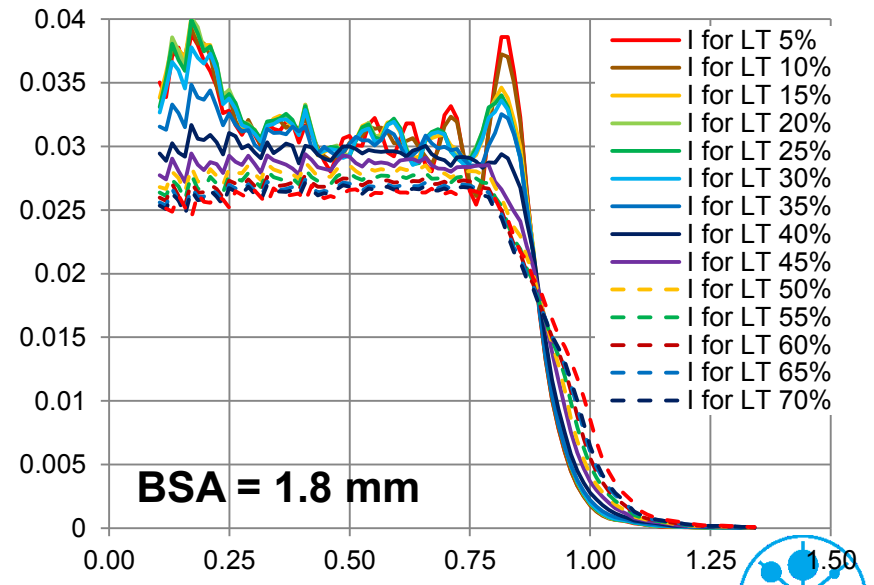
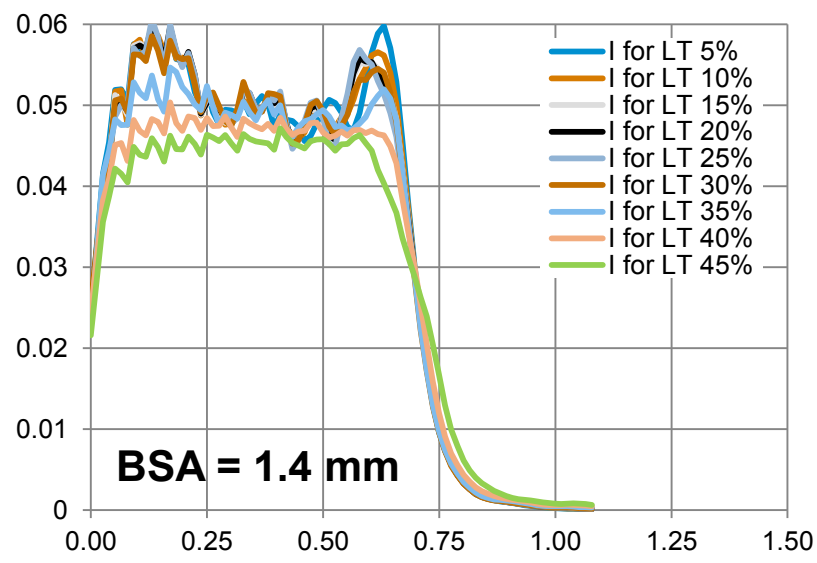
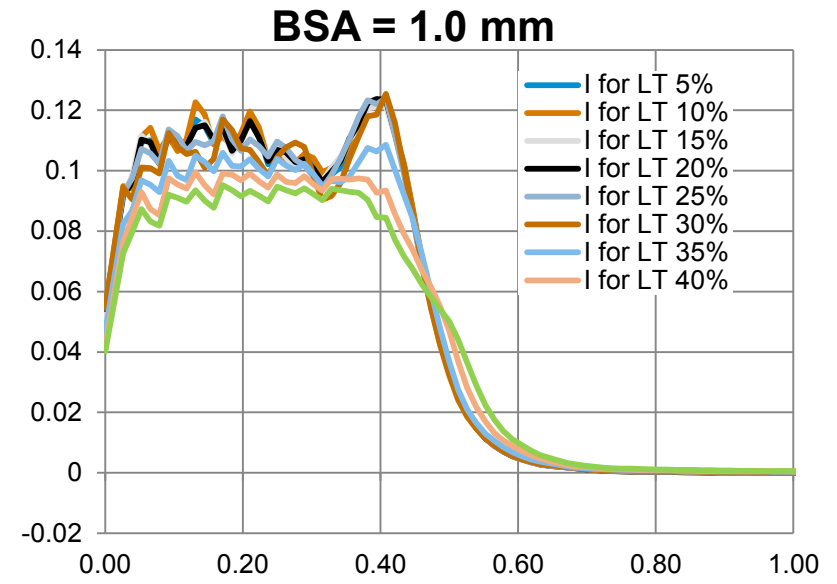
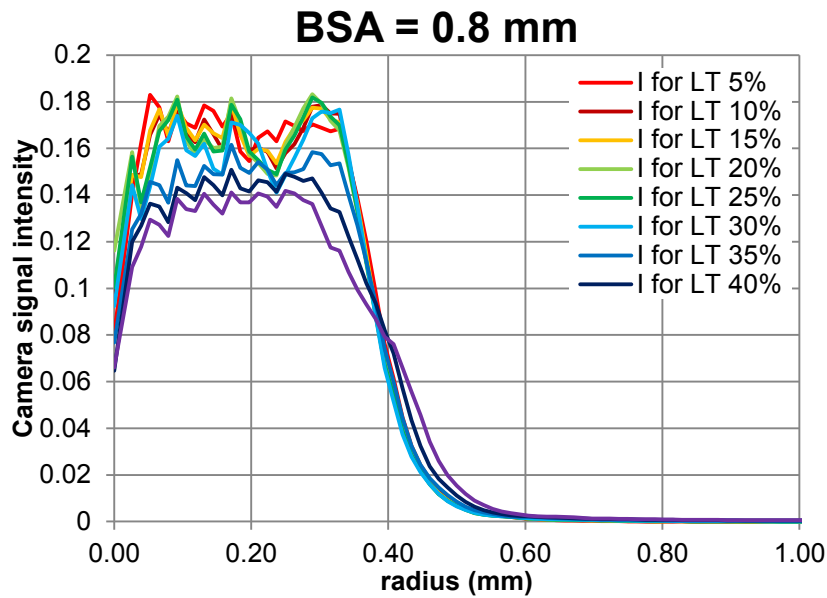


> BACKUP

> SLIDES



Laser radial profiles extracted from measured transverse distributions (integrated over ϕ) for various beam diameters



The output from postpro shows that XYrms for core+halo distribution is very close to that for the uniform distribution

```
Graphic1
-----
200000 particles from file c+h_Rc=034_sg=013.ini

Cathode located at:          z =      0.000      m
Particles taken into account N =    200000
total charge                 Q =   -0.1624      nC
horizontal beam position     x =   5.0998E-04 mm
vertical beam position       y =   1.2130E-03 mm
longitudinal beam position   z =      0.000      m
horizontal beam size         sig x =   0.1984      mm
vertical beam size           sig y =   0.1989      mm
longitudinal beam size       sig z =      0.000      mm
total emission time          t =   1.3755E-02 ns
rms emission time            sig t =   1.5038E-03 ns
average kinetic energy       E =   5.4988E-07 MeV
energy spread                 dE =   1.6919E-06 keV
average momentum             P =   7.4965E-04 MeV/c
transverse beam emittance    eps x =   0.1679      pi mrad mm
correlated divergence        cor x =   -8.178      mrad
transverse beam emittance    eps y =   0.1686      pi mrad mm
correlated divergence        cor y =   0.1764      mrad
longitudinal beam emittance  eps z =      0.000      pi keV mm
correlated energy spread     cor z =      0.000      keV
emittance ratio eps y/eps x  =   0.9956

trace space emittance        eps x =   45.84      pi mrad mm
trace space emittance        eps y =   48.99      pi mrad mm
Reduced emittances:
hor. emittance minus z correlation: =   0.1679      pi mrad mm
hor. emittance minus z & E correlation: =   0.1609      pi mrad mm
ver. emittance minus z correlation: =   0.1686      pi mrad mm
ver. emittance minus z & E correlation: =   0.1509      pi mrad mm
long. emittance minus 2nd order corr.: =   0.000      pi keV mm
long. emittance minus 2nd & 3rd order corr.: =   0.000      pi keV mm

Particle Statistics:

Total number of particles on stack =   200000
Electrons (total)                 =   200000
particles at the cathode           =   200000
active particles                    =   0
passive particles (lost out of bunch) =   0
```



Size (slide by M. Gross)

BSA size [mm]*	Cathode camera			VC2			Size ratios Cathode/VC2			BSA/VC2 xyRMS
	xRMS	yRMS	xyRMS	xRMS	yRMS	xyRMS	x	y	xy	
0.08	0.066	0.065	0.065	0.064	0.063	0.063	1.03	1.03	1.03	1.26
0.16	0.056	0.068	0.062	0.055	0.062	0.058	1.02	1.10	1.06	2.74
0.22	0.072	0.085	0.078	0.067	0.078	0.072	1.07	1.09	1.08	3.04
0.33	0.095	0.107	0.101	0.091	0.106	0.098	1.04	1.01	1.03	3.36
0.75	0.202	0.21	0.206	0.203	0.205	0.204	1.00	1.02	1.01	3.68
1.13	0.3	0.301	0.300	0.306	0.294	0.300	0.98	1.02	1.00	3.77
1.52	0.413	0.402	0.407	0.409	0.4	0.404	1.01	1.01	1.01	3.76
1.98	0.511	0.499	0.505	0.504	0.493	0.498	1.01	1.01	1.01	3.97
2.37	0.624	0.604	0.614	0.618	0.588	0.603	1.01	1.03	1.02	3.93
3.5	0.843	0.817	0.830	0.836	0.791	0.813	1.01	1.03	1.02	4.30

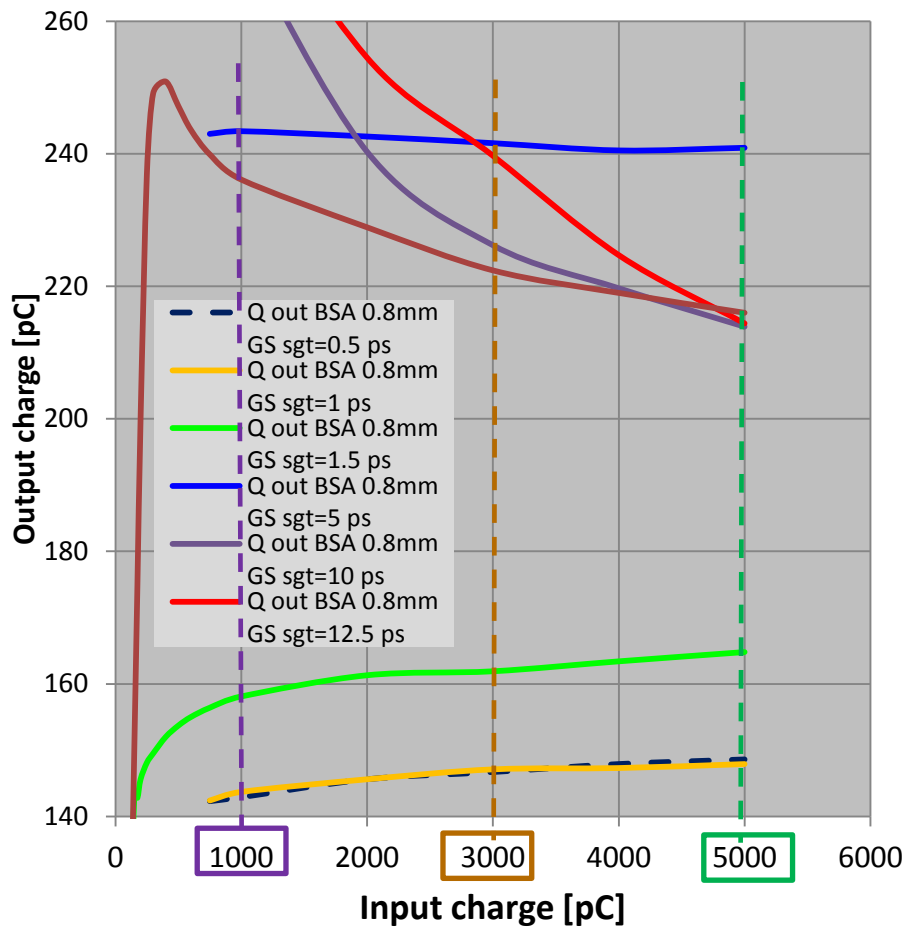
Average: 1.02 1.04 1.03

- > Laser beam a little bit bigger on photocathode (1 to 3%)
- > *BSA (calibration?) – Ratio BSA size to xyRMS about 4 for flat tops, reduced for smaller sizes
 - No flat top for small BSA

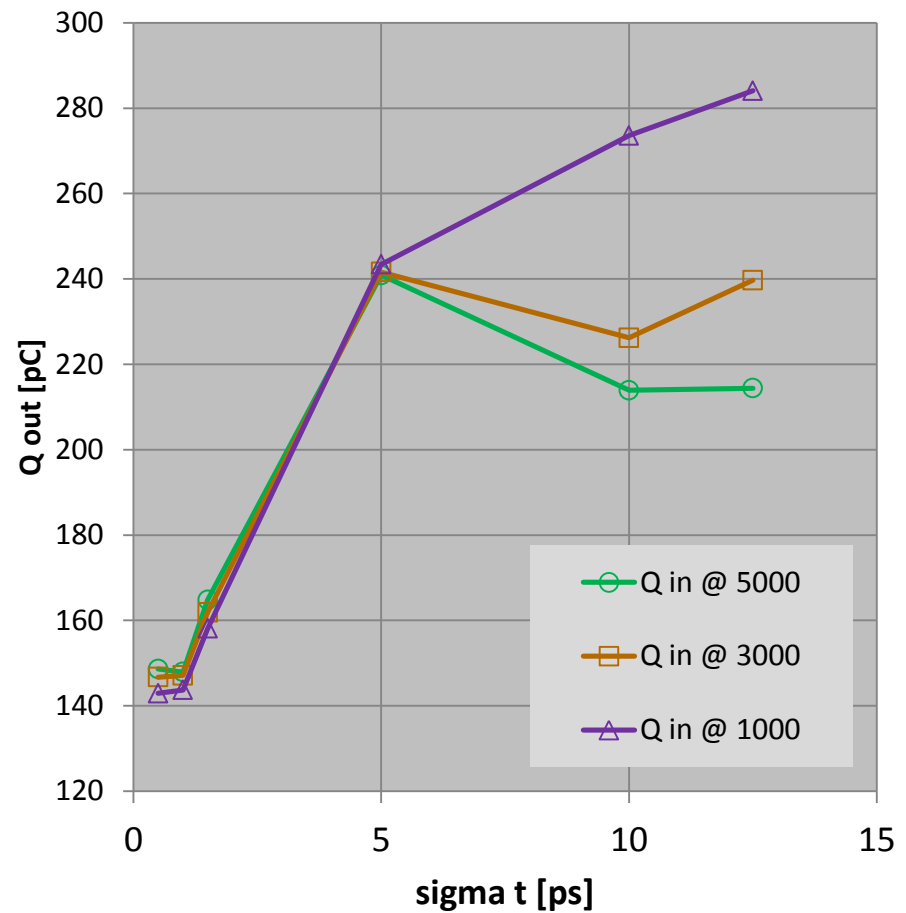


The slope of the output charge depends on the initial bunch length

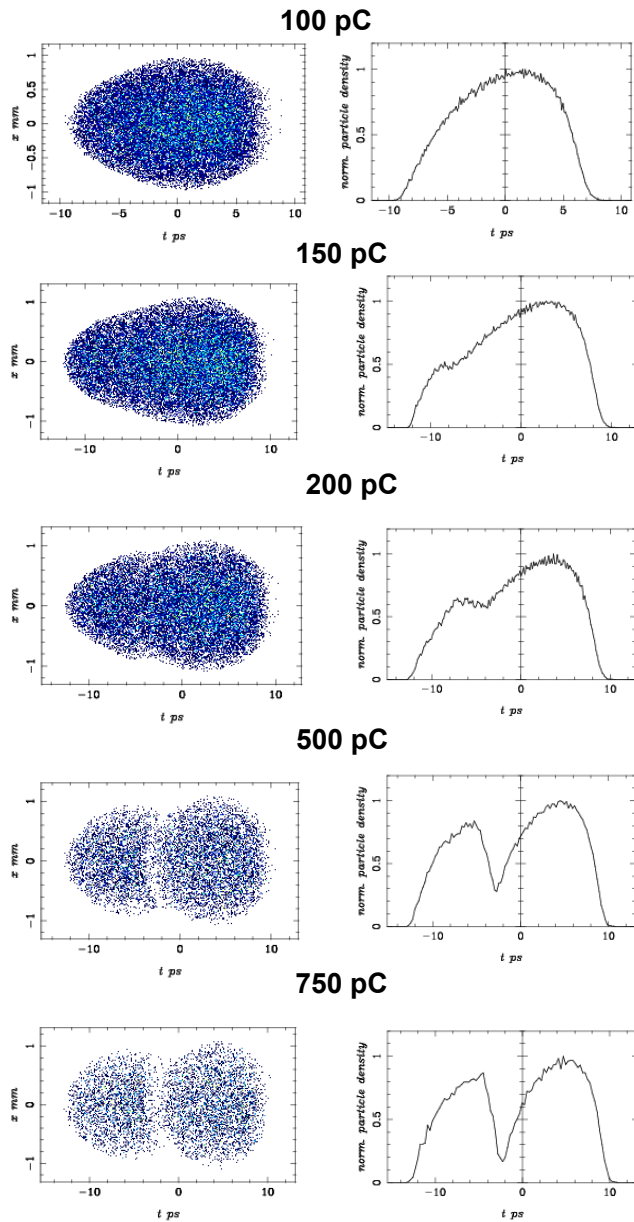
Output charge vs injected charge for Gaussian temporal profiles with varying pulse length



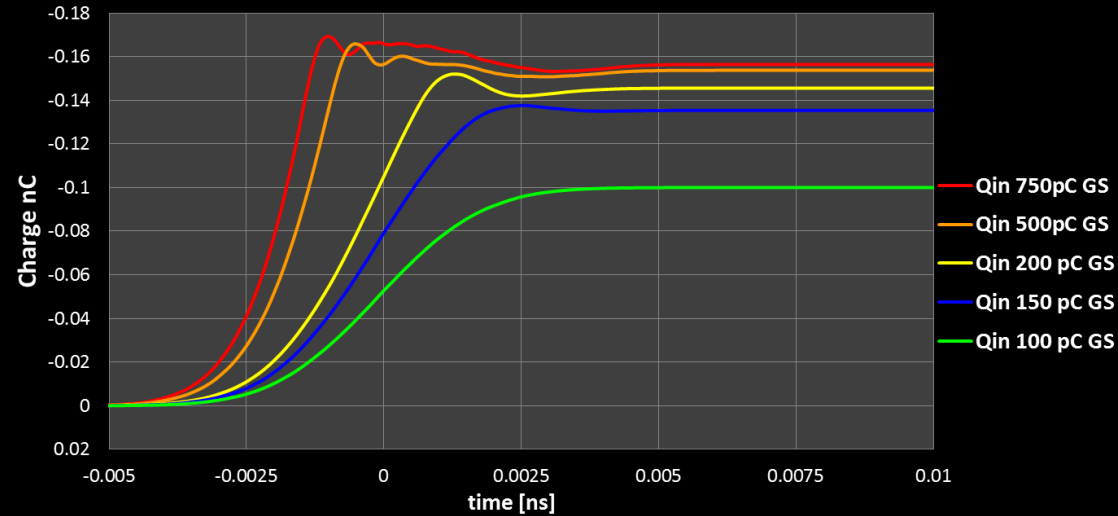
Output charge evolution with time for indicated set of injected charges



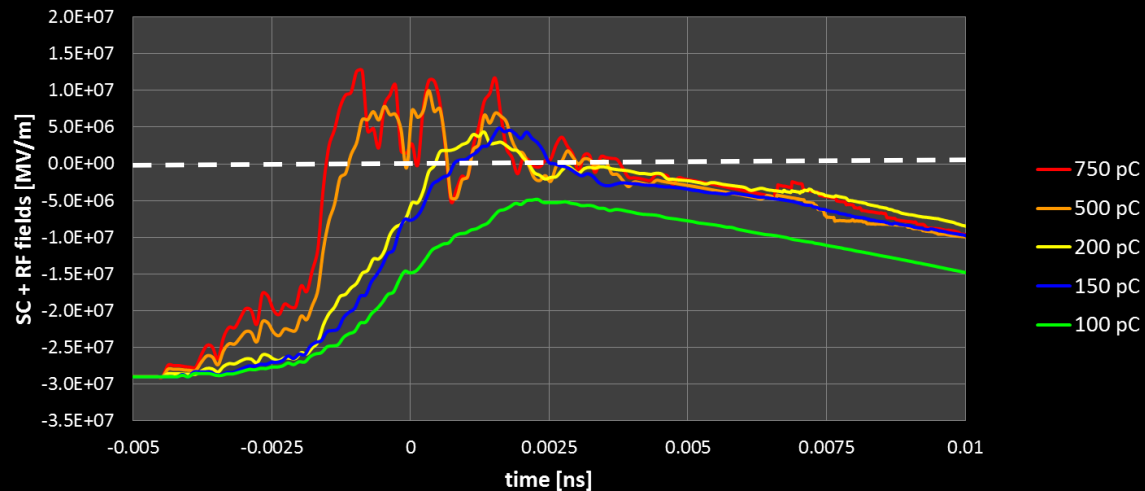
More charge can be extracted as the injected current is increased due to the formation of the virtual cathode



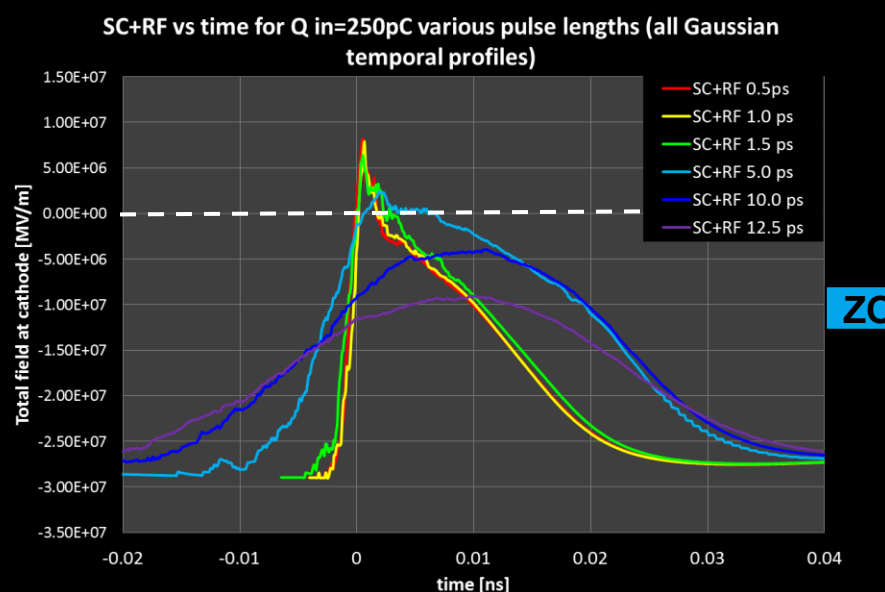
Output charge vs t for Gaussian temporal with $\text{sig}_t=1.5$ ps for various injected charges



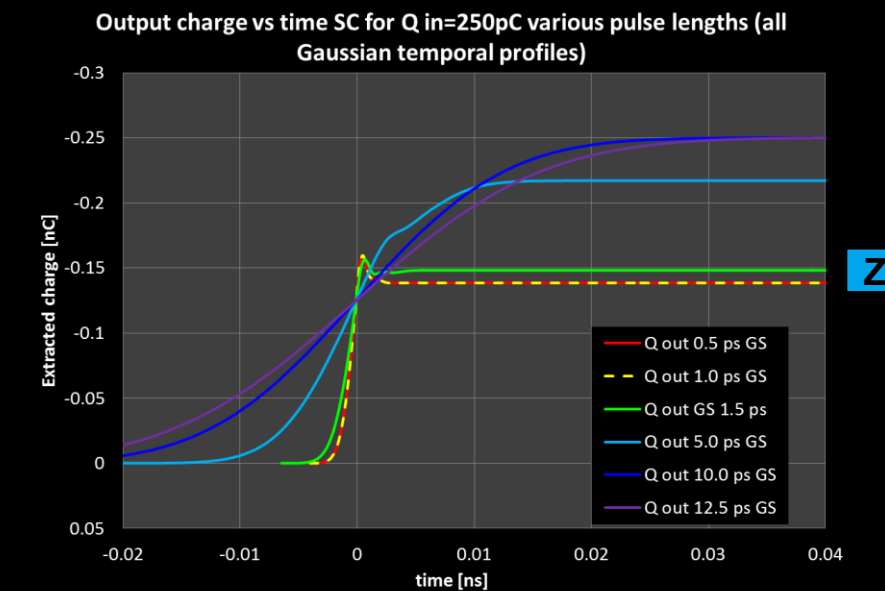
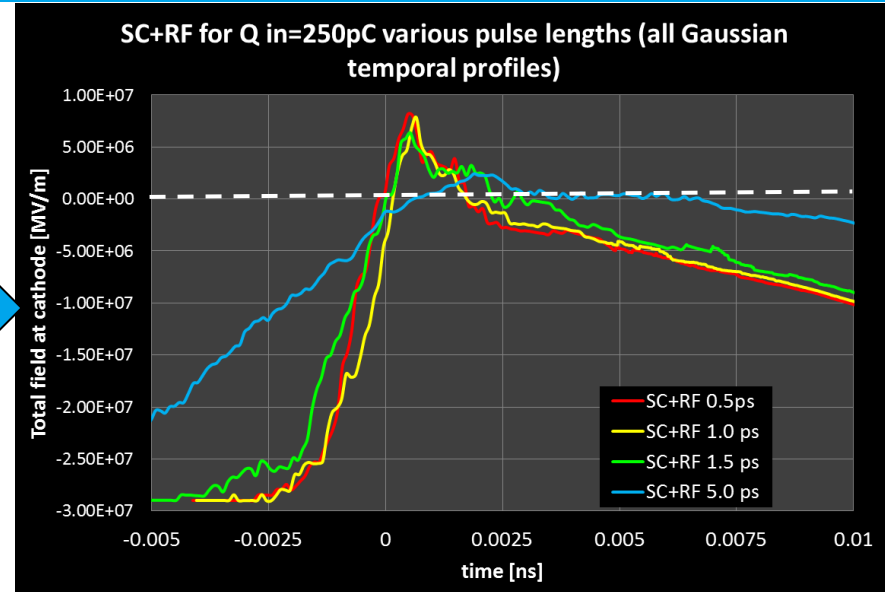
Cathode Total Electric field (SC+RF) vs time for Gaussian temporal $\text{sig}_t=1.5$ ps for various injected charges



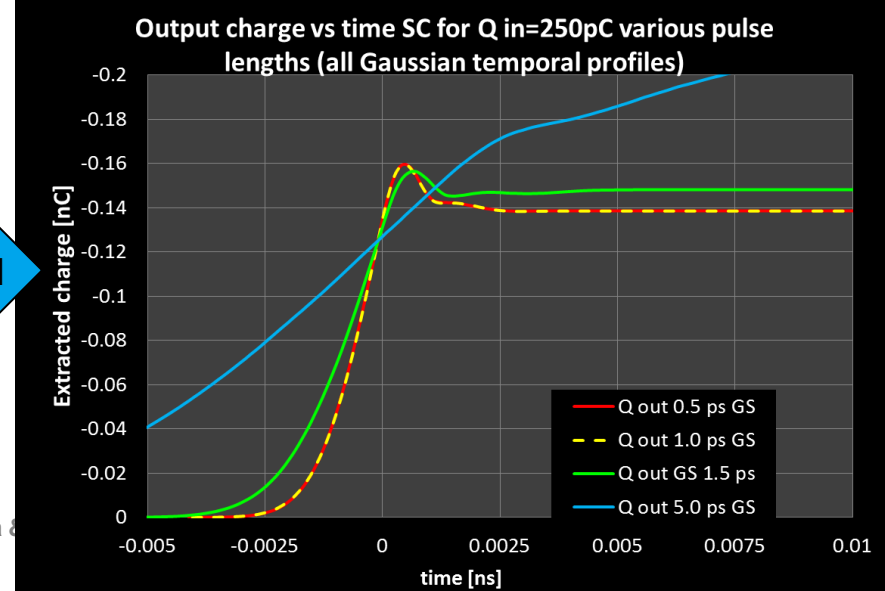
The formation of the virtual cathode is characterized by oscillations in the space charge field at the surface. As the particles at the head of the bunch move away from the surface, the charge density drops increasing the field at the cathode which in turn allows for more charge to be extracted. The process repeats in an oscillatory fashion.



ZOOM IN



ZOOM IN



In addition, QE scans for old and new cathodes are drastically different.

- How to reproduce these measurements in simulations if machine parameters are the same for both?
- $P_{\text{gun}} = 1.5 \text{ MW}$ at $\Phi_0\text{-}90$ and $\text{BSA} = 1.8 \text{ mm}$

