

Calculations of Transition Radiation from an Relativistic Electron Beam

Outline

- ▶ Introduction
- ▶ TR from a Point Charge
- ▶ TR from an Electron Bunch
- ▶ Algorithms of Numerical Calculations
- ▶ Example Results of Numerical Calculations
- ▶ Summary and Outlook
- ▶ References

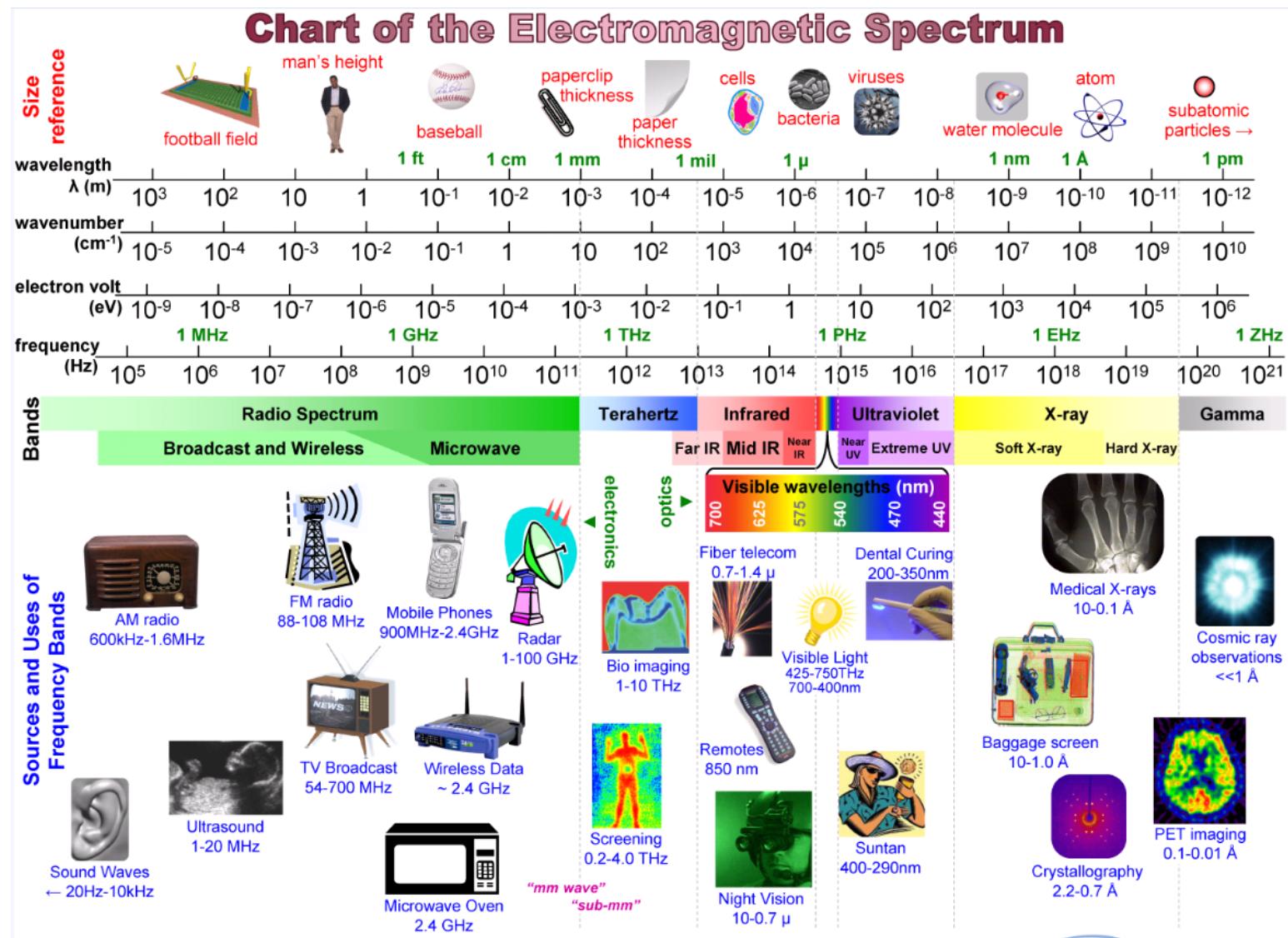
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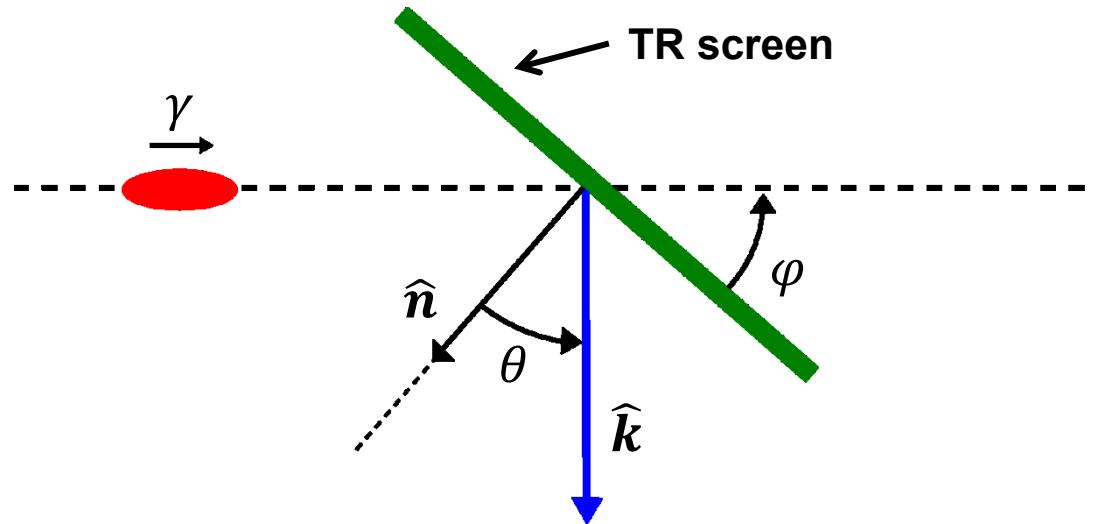
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$$\lambda = 3 \times 10^8 / \text{freq} = 1 / (\omega n * 100) = 1.24 \times 10^{-6} / \text{eV}$$

Calculations of Transition Radiation from an Relativistic Electron Beam
Prach Boonpornprasert, PITZ Physics Seminar, 23.04.2015

Introduction: Transition Radiation (TR)

TR is emitted when a relativistic charged particle crosses the boundary between two media of different dielectric properties.



- γ is Lorentz factor of electron.
- \hat{n} is normal vector of the TR screen.
- \hat{k} is wave vector of outgoing radiation.
- θ is inclination angle between \hat{k} and \hat{n} .
- φ is angle between screen plane and electron beam axis.

TR from a Point Charge

- TR spectral energy density can be calculated from Ginzburg-Frank Formula:

$$\frac{d^2 U_{GF}}{d\omega d\Omega} = \frac{e^2}{4\pi^3 \varepsilon_0 c} \frac{\beta^2 \sin^2 \theta}{(1 - \beta^2 \cos^2 \theta)^2}$$

where U_{GF} is radiation energy of backward radiation from Ginzburg-Frank formula.
 ω is angular frequency of radiation
 Ω is solid angle that subtends radiation
 ε_0 is electric constant.
 β is ratio between electron velocity and speed of light.

- The formula is valid in the following conditions:

- Normal Incident ($\varphi = \pi/2$)
- A single relativistic electron
- Infinite metallic screen of perfect reflectivity
- Far-field

TR from a Point Charge (2)

► Generalized Ginzburg-Frank Formula:

$$\frac{d^2 U_{\text{disk}}}{d\omega d\Omega} = \frac{d^2 U_{\text{GF}}}{d\omega d\Omega} [1 - T(\gamma, \omega a, \theta)]^2$$

where U_{disk} is radiation energy of backward radiation from generalized Ginzburg-Frank formula.

a is radius of the circular TR screen.

$$T(\gamma, \omega a, \theta) = \frac{\omega a}{c\beta\gamma} J_0 \left(\frac{\omega a \sin \theta}{c} \right) K_1 \left(\frac{\omega a}{c\beta\gamma} \right) + \frac{\omega a \sin \theta}{c\beta^2\gamma^2 \sin \theta} J_1 \left(\frac{\omega a \sin \theta}{c} \right) K_0 \left(\frac{\omega a}{c\beta\gamma} \right)$$

J_n, K_n are Bessel functions.

► The formula is valid in the following conditions:

- Normal Incident ($\varphi = \pi/2$)
- A single relativistic electron
- Finite circular TR screen with radius a
- Far-field

TR from an Electron Bunch

- TR spectral energy density from an electron bunch can be calculated from the following formula:

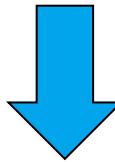
$$\frac{d^2 U_{\text{bunch}}}{d\omega d\Omega} = \frac{d^2 U_{\text{GF}}}{d\omega d\Omega} N^2 |F_{\text{long}}(\omega)|^2 \cdot \left[\frac{2c}{\omega r_b \sin \theta} J_0\left(\frac{\omega r_b \sin \theta}{c}\right) - \frac{2c\beta\gamma}{\omega r_b} I_0\left(\frac{\omega r_b}{c\beta\gamma}\right) T(\gamma, \omega a, \theta) \right]^2$$

where U_{disk} is radiation energy of backward radiation from generalized Ginzburg-Frank formula.
 N is total number of electrons in a bunch.
 r_b is transverse radius of electron bunch.
 $\rho_{\text{long}}(t)$ is longitudinal particle density distribution of electron bunch.
 $F_{\text{long}}(\omega) = \int_{-\infty}^{+\infty} \rho_{\text{long}}(t) e^{-i\omega t} dt$ ← “Form Factor”
 I_n is Bessel function.

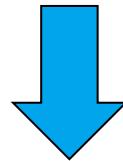
- The formula is valid in the following conditions:

- Normal Incident ($\varphi = \pi/2$)
- A relativistic electron bunch with r_b and $\rho_{\text{long}}(t)$
- Finite circular TR screen with radius a
- Far-field

Input electron beam file
(ASTRA, CSRTrack, etc.)



CalCTR code
(developed on MATLAB)



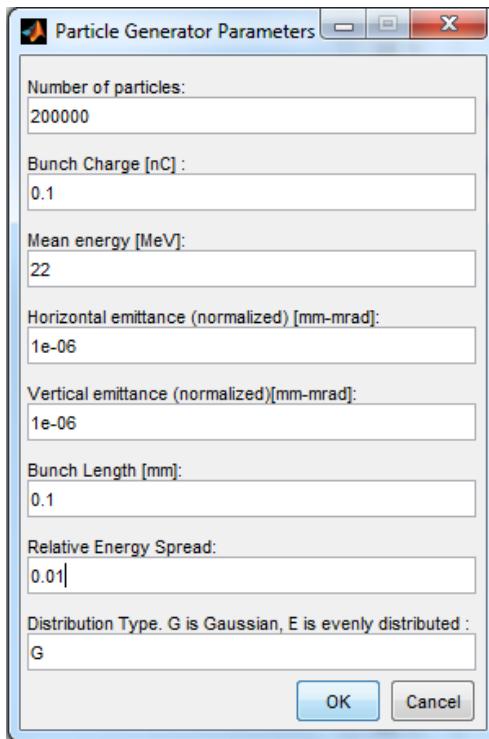
Output CTR radiation

**Will be presented
with more details later**

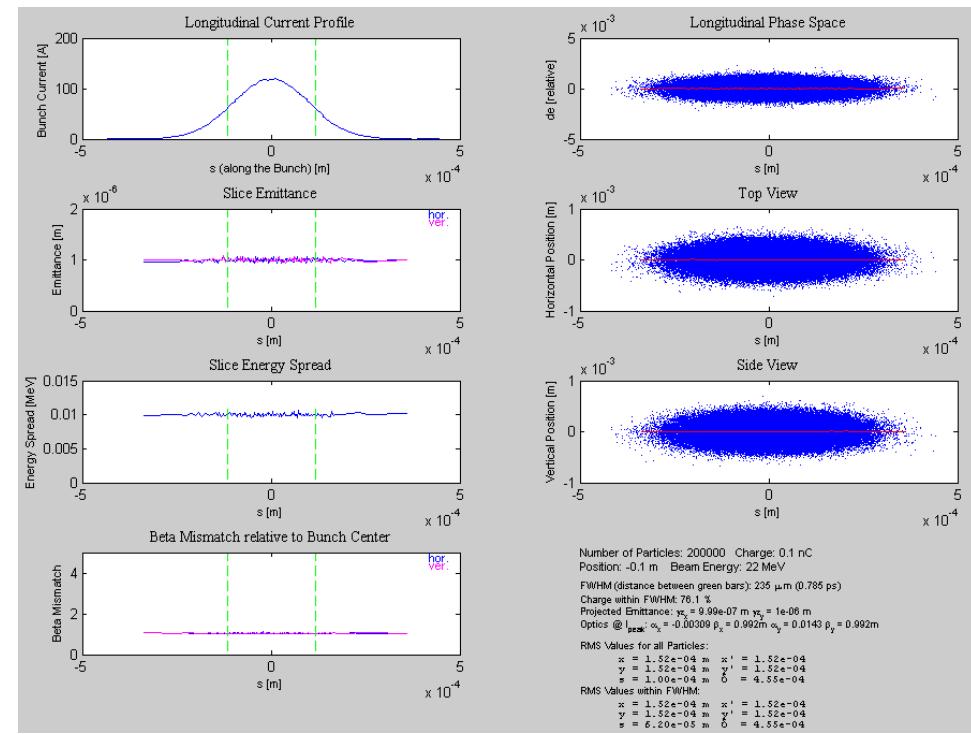
Preparation of Input Beams

- Generated model Gaussian beams from PSViewer software with various parameters.

The Input Box

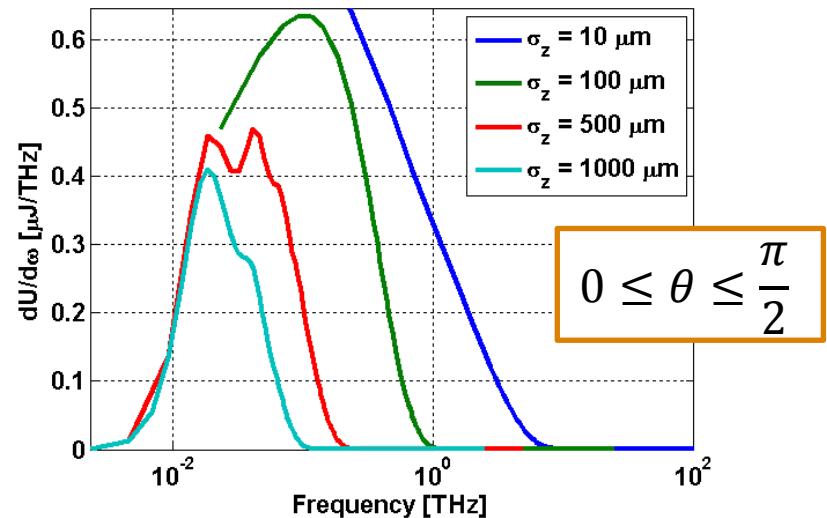
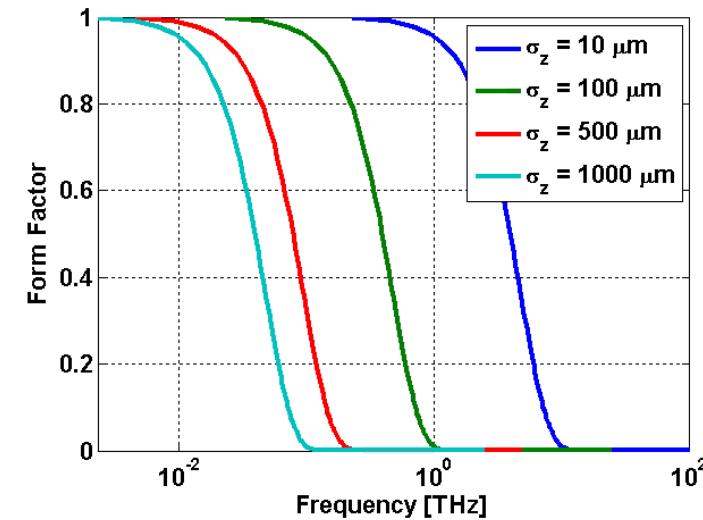
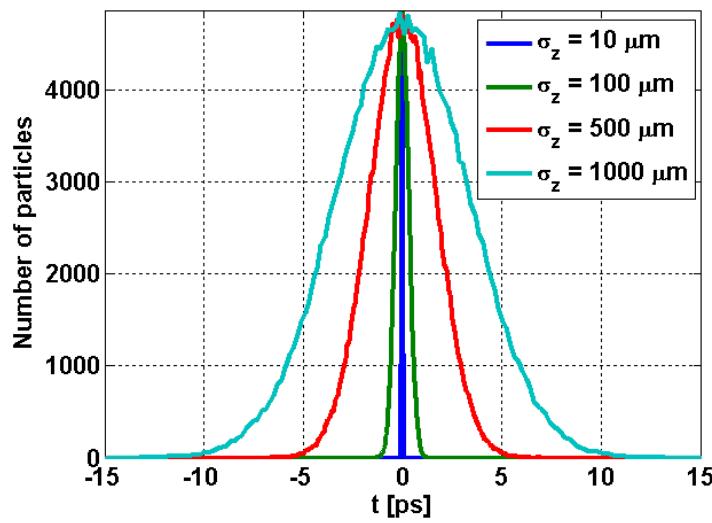


Example of Generated Beam Profile



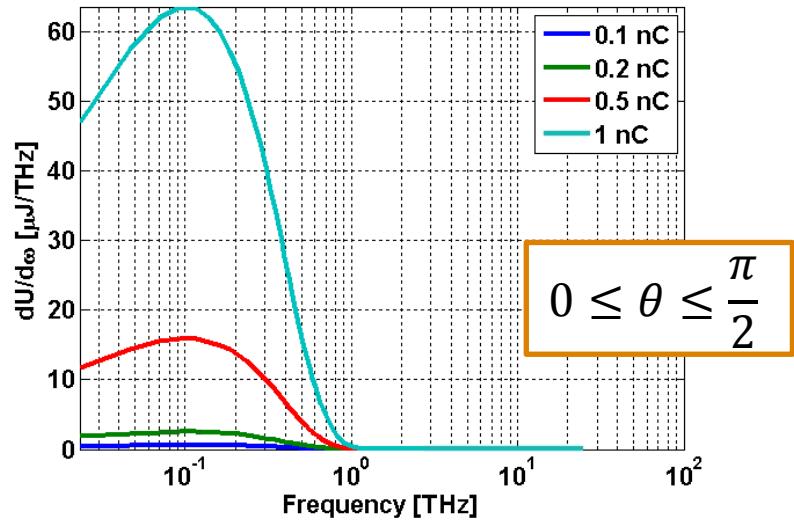
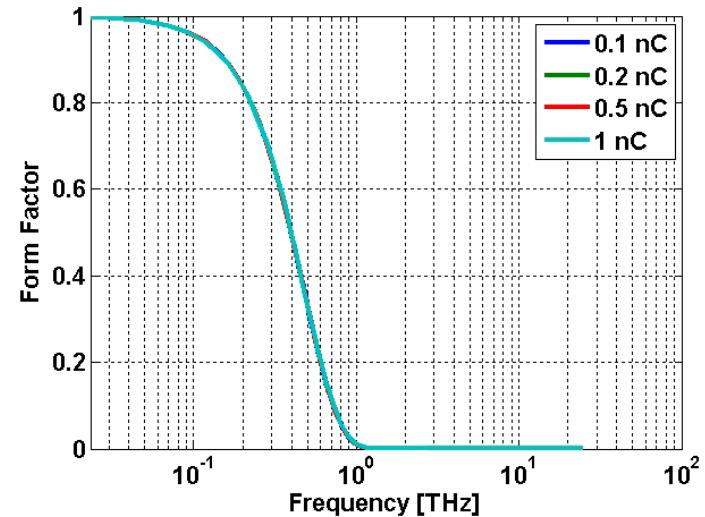
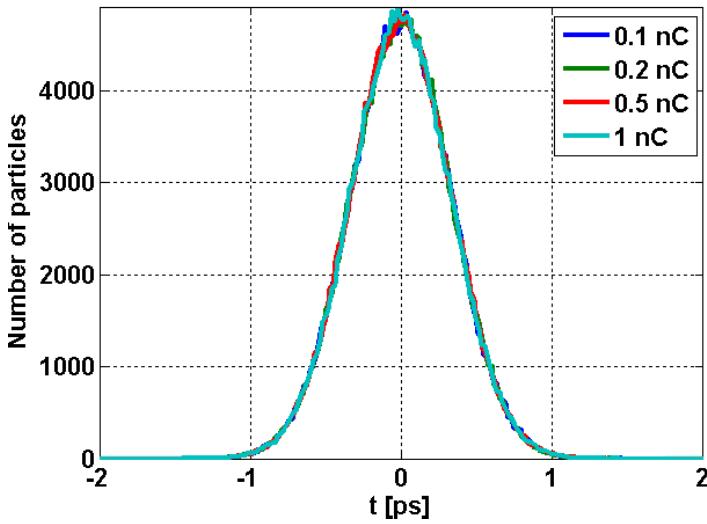
Effects of Bunch Lengths

Input Parameter	Value
Number of particles	200k
Bunch charge	0.1 nC
Mean energy	22 MeV
Relative energy spread	0.01
Tr. Emittance	1 mm mrad
Bunch rms length	10,50,100,1000 μm
Tr. rms size	0.152 mm



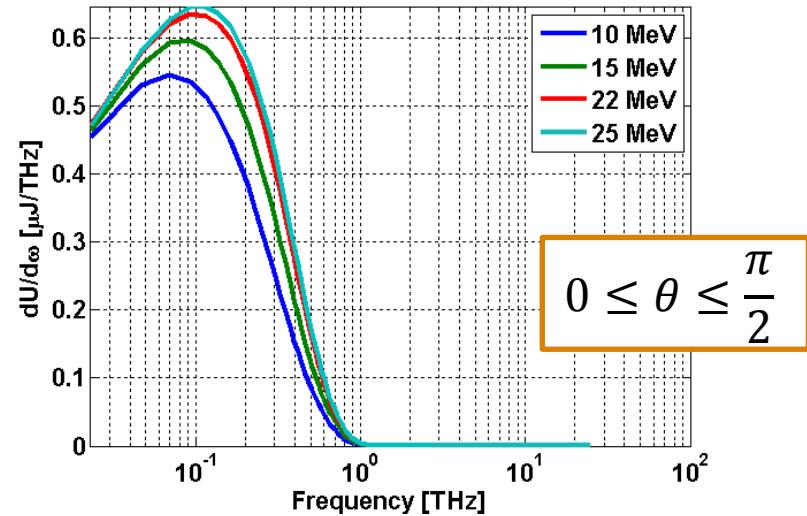
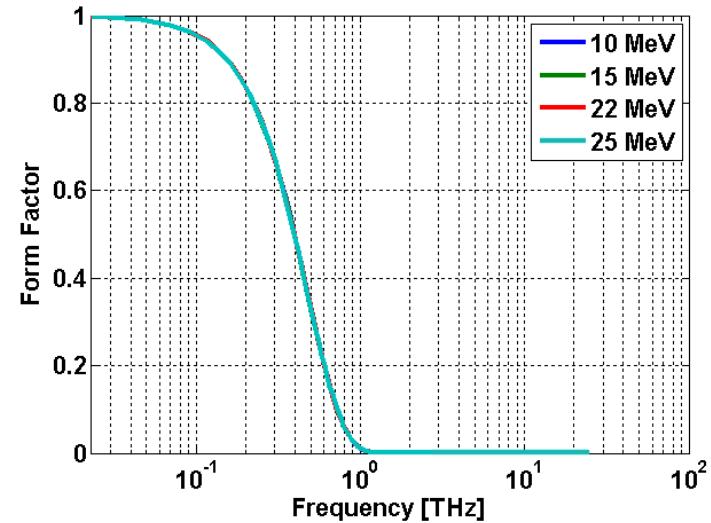
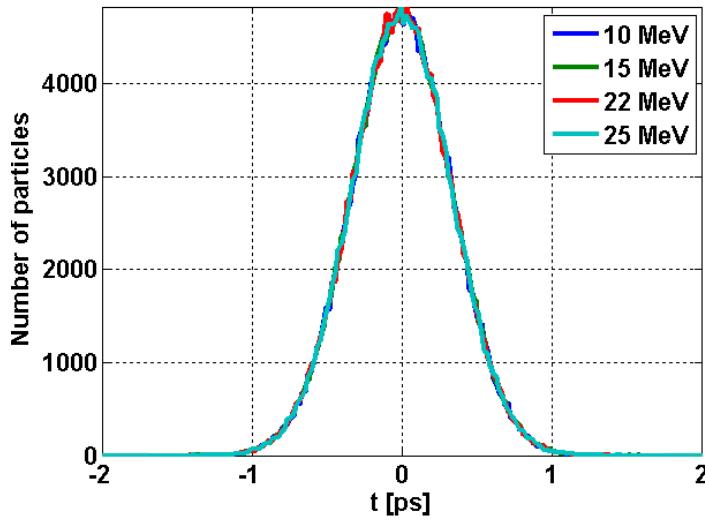
Effects of Peak Currents

Input Parameter	Value
Number of particles	200k
Bunch charge	0.1, 0.2, 0.5, 1 nC
Mean energy	22 MeV
Relative energy spread	0.01
Tr. Emittance	1 mm mrad
Bunch rms length	100 μm
Tr. rms size	0.152 mm



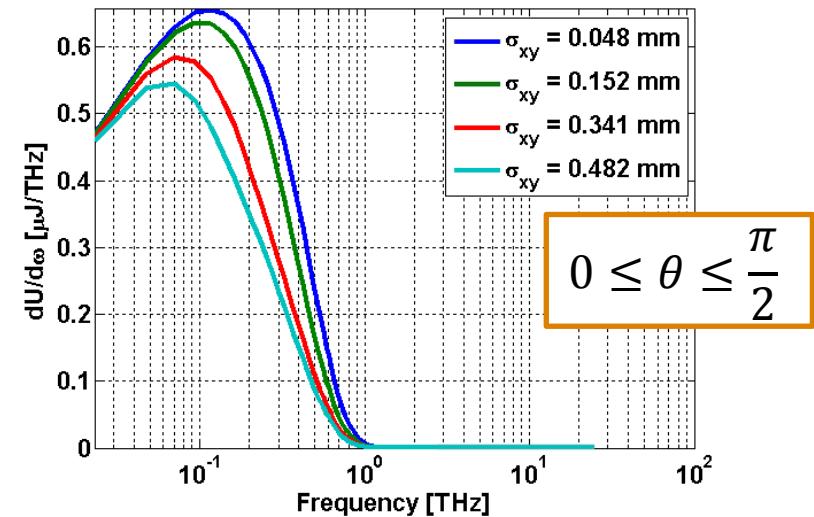
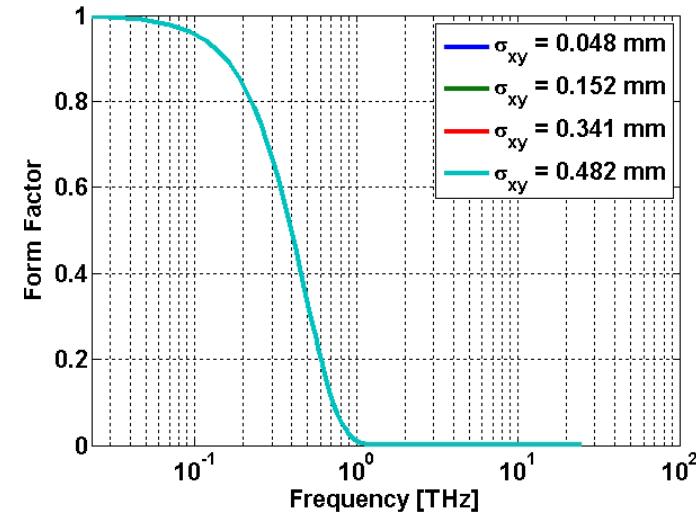
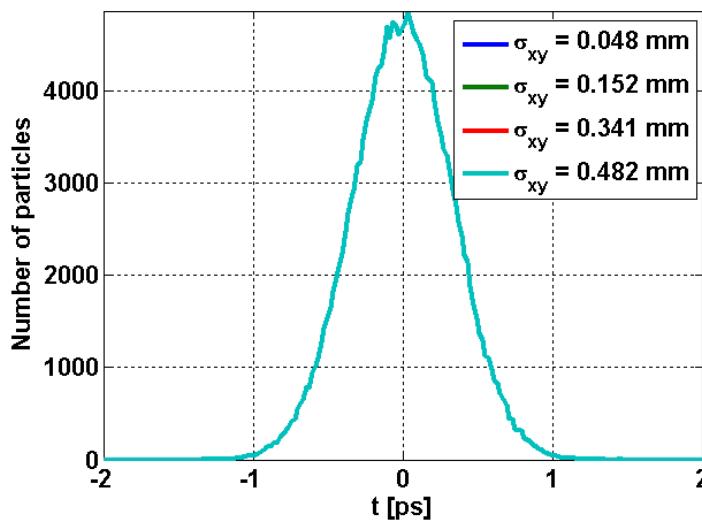
Effects of Beam Energies

Input Parameter	Value
Number of particles	200k
Bunch charge	0.1 nC
Mean energy	10, 15, 22, 25 MeV
Relative energy spread	0.01
Tr. Emittance	1 mm mrad
Bunch rms length	100 μm
Tr. rms size	0.152 mm



Effects of Beam Sizes

Input Parameter	Value
Number of particles	200k
Bunch charge	0.1 nC
Mean energy	22 MeV
Relative energy spread	0.01
Tr. Emittance	1 mm mrad
Bunch rms length	100 μ m
Tr. rms size	0.048, 0.152, 0.341, 0.482 mm



Summary and Outlook

Summary

- ▶ Beta version of CalCTR has been developed.
- ▶ Effects of electron beam parameters to the CTR energy spectrum density were studied.

Outlook

- ▶ Implement in CalCTR:
 - sliced energy
 - slice beam size
 - oblique target screen
 - near-field calculations
 - electric field of THz pulse
- ▶ Simulation of velocity bunching (PPS beginning of May)
- ▶ Calculation of TR from PITZ Beam (PPS beginning of May)
- ▶ Consideration of TR diagnostics (PPS end of May)

References

- ▶ S.Casalbuoni et al., TESLA Report 2005-15,
http://tesla.desy.de/new_pages/TESLA_Reports/2005/pdf_files/tesla2005-15.pdf
- ▶ S.Casalbuoni et al., TESLA-FEL 2006-04,
http://flash.desy.de/sites2009/site_vuvfel/content/e403/e1642/e1132/e1129/infoboxContent1745/fel2006-04.pdf
- ▶ M. Schwarz et al., IPAC2012, MOPPP003,
<http://accelconf.web.cern.ch/AccelConf/IPAC2012/papers/moppp003.pdf>
- ▶ S. Naknaimueang et al., IPAC2011, TUPO007,
<http://accelconf.web.cern.ch/AccelConf/IPAC2011/papers/tupo007.pdf>
- ▶ C. Settakorn, SLAC-R-576,
<http://www.slac.stanford.edu/cgi-wrap/getdoc/slac-r-576.pdf>
- ▶ K.N.Ricci and T. I.Smith, Phys. Rev. ST Accel. Beams 3, 032801
<http://journals.aps.org/prstab/abstract/10.1103/PhysRevSTAB.3.032801>
- ▶ O.Grimm and P.Schmüser, TESLA FEL 2006-03,
http://flash.desy.de/sites2009/site_vuvfel/content/e403/e1642/e1132/e1129/infoboxContent1498/fel2006-03.pdf

Thanks for your attentions!